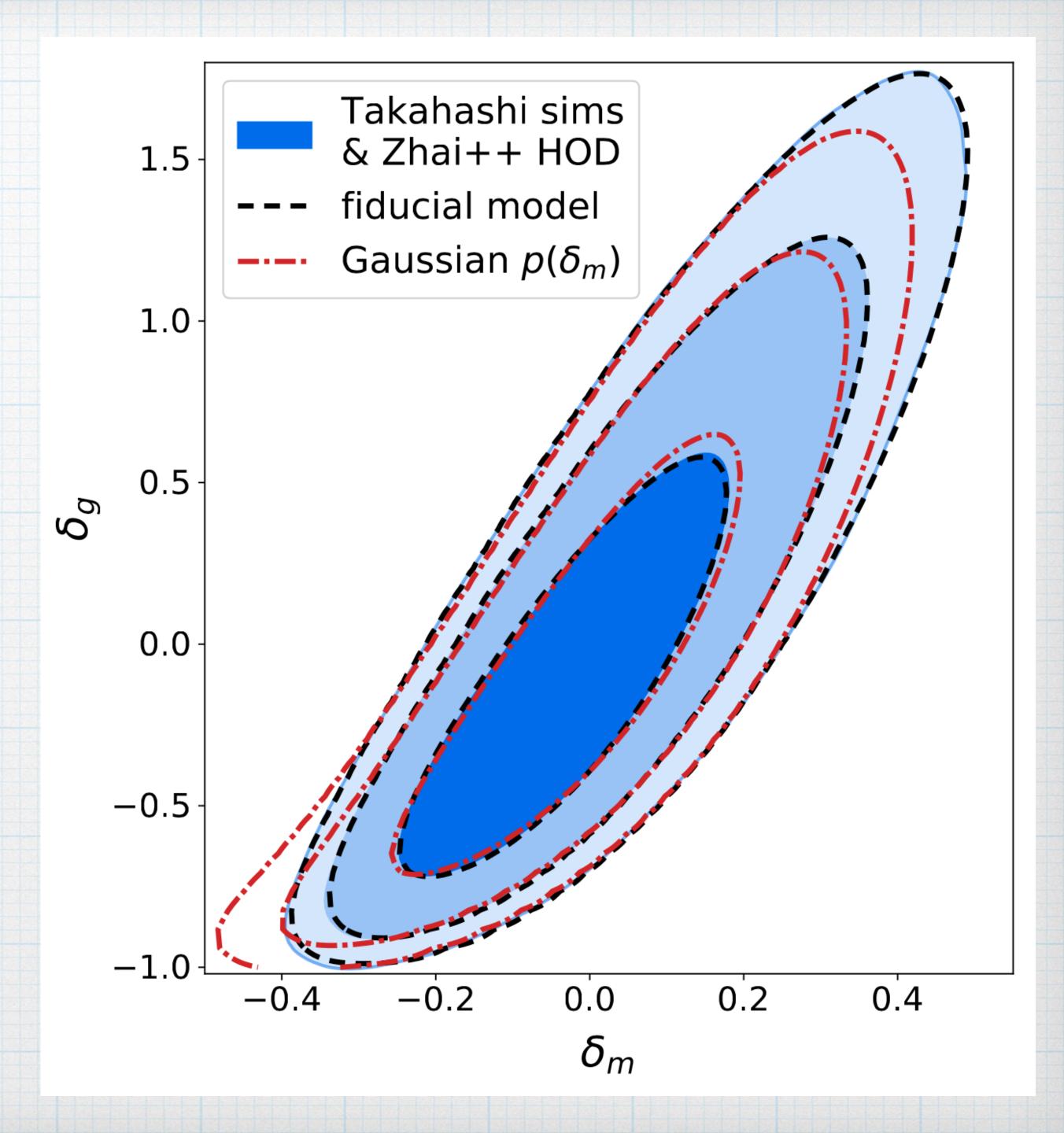
### Analysing the full shape of the PDF of cosmic density fluctuations

Oliver Friedrich Newton-Kavli Junior Fellow

work with Cora Uhlemann, Anik Halder, Aoife Boyle, Dylan Britt, Sandrine Codis, Daniel Gruen, ChangHoon Hahn, Tobias Baldauf, Francisco Villaescusa-Navarro, Marc Manera, Takahiro Nishimichi and more!



But first... You may know me from my covariance work.



### But first... You may know me from my covariance work. $\rightarrow$ I'm retiring from this subject



### But first... You may know me from my covariance work. $\rightarrow$ I'm retiring from this subject

- $\rightarrow$  Last shout-out to my work on that:
- Friedrich et al. (2016) Internal covariance estimation
- Friedrich & Eifler (2018) Have a bit of analytic knowledge of your covariance?
- Friedrich et al. (2021a) "Instruction manual" on covariance modelling, estimation and validation for multi-probe 2-point functions
- Percival, Friedrich, Selentin, Heavens (2021) Taking into account covariance uncertainties in parameter constraints



2-point statistics: Variance as a function of scale

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 $\langle \delta_{\mathbf{k}} \delta_{\mathbf{q}} \rangle \sim P(k) \ \delta_D(\mathbf{k} + \mathbf{q})$ 

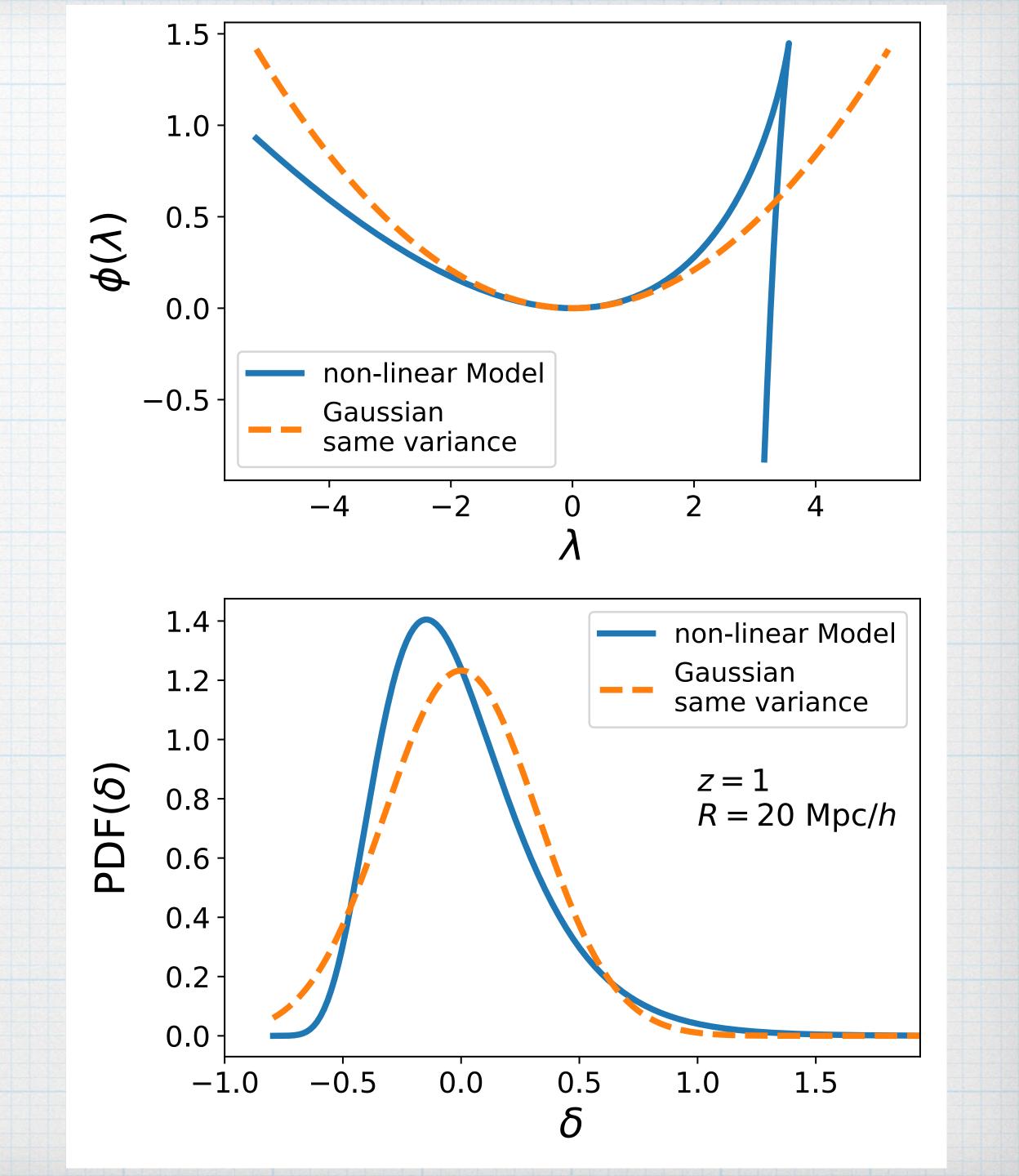
Cumulant generating function (CGF): All moments at one scale

 $\varphi_R(\lambda) = \langle e^{\lambda \delta_R} \rangle \equiv \sum \langle \delta_R^n \rangle_c \frac{\lambda^n}{n!}$ 

n

 PDF and CGF related by Laplace transform

 $e^{\varphi_R(\lambda)} = d\delta_R p(\delta_R) e^{\lambda\delta_R}$ 1 1



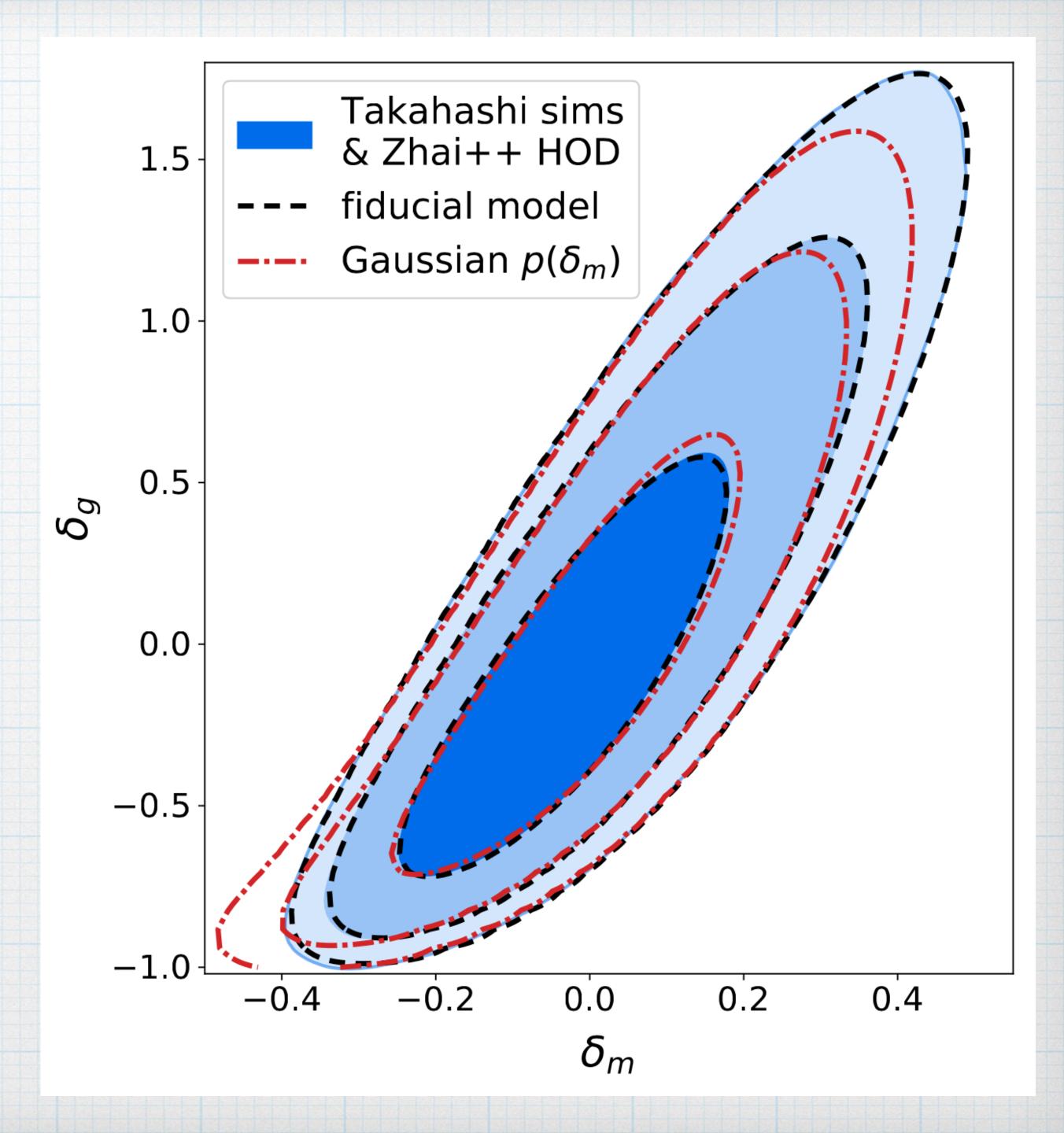
.... now including galaxies.

"3 x 2-point analyses": compressing joint PDF of galaxy and matter density to 3 numbers!

cosmic shear ~ variance of matter density gal. clustering ~ variance of galaxy density gal.-gal. lensing ~ covariance of the two

 $\rightarrow$  Can we do better?

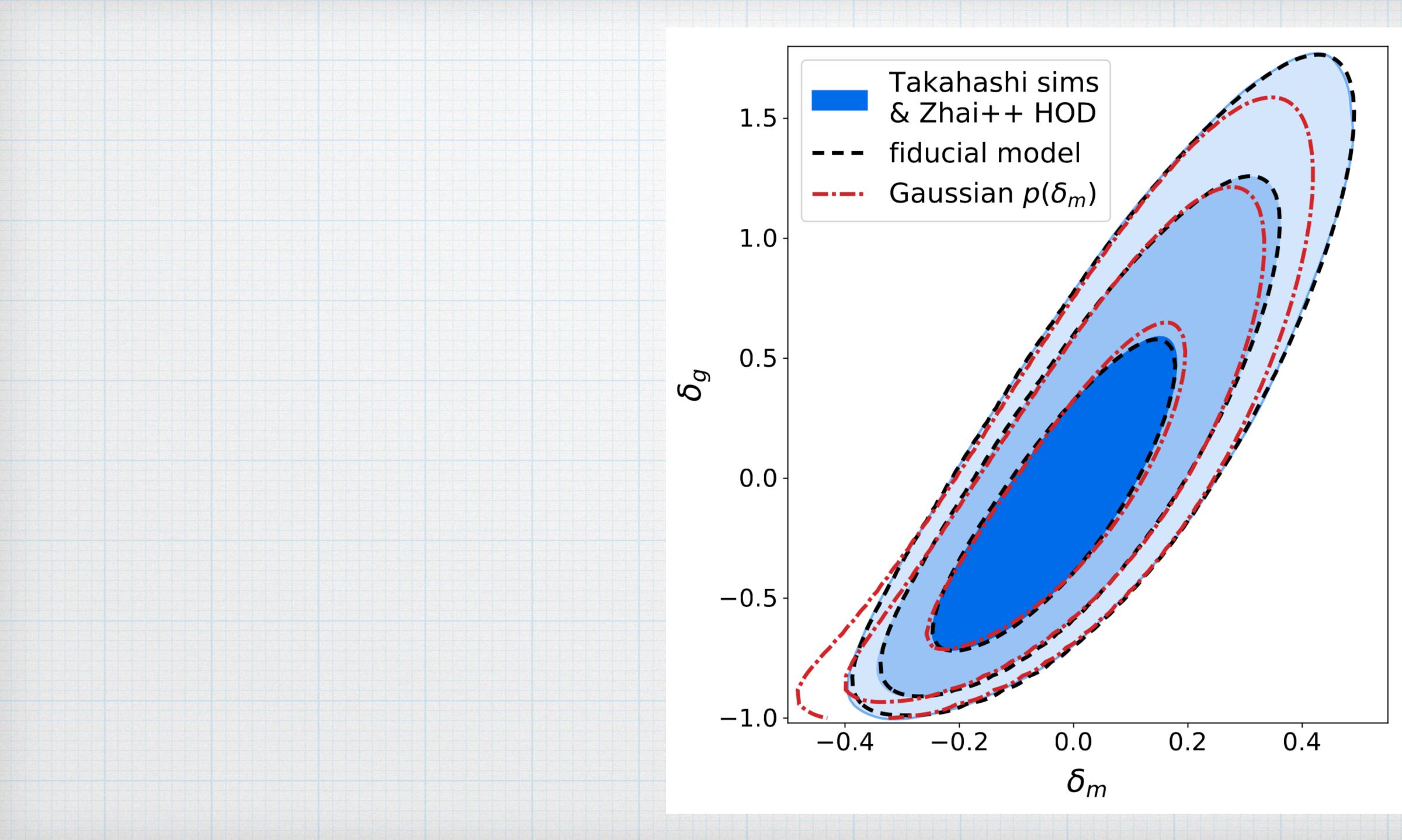
 $\rightarrow$  Is it worth it to do better?



#### Recent work:

- Friedrich et al. (2020) uncertainty
- Uhlemann, Friedrich et al. (2020), Boyle, Uhlemann, Friedrich et al. (2021) **2-point function** Strongly improved constraints on neutrino mass and DE equation of state
- Friedrich et al. (2021b)
  - non-Poisson shot-noise
  - -> 2D language appropriate for photometric data







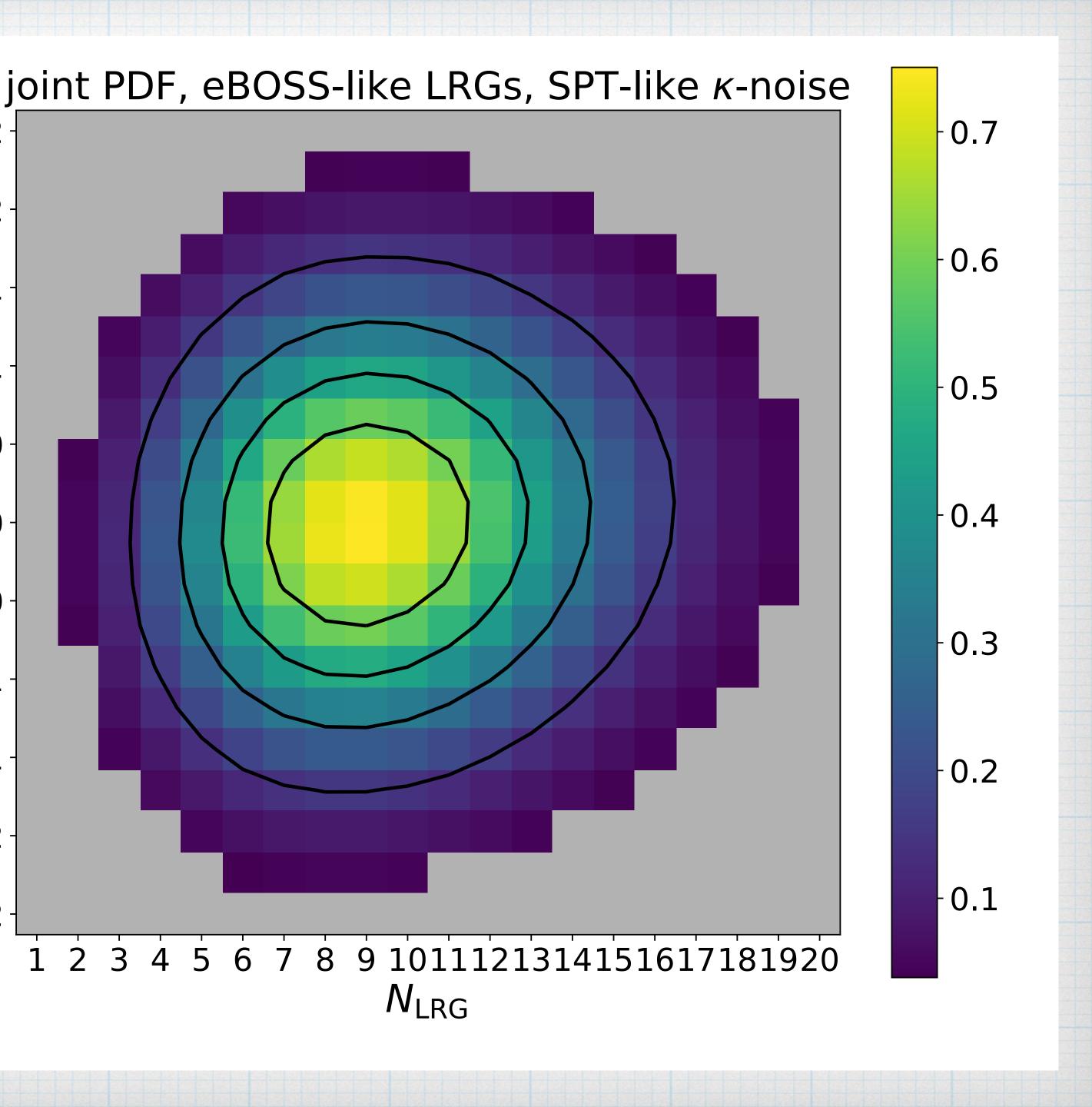
Moving to real data: Friedrich ++ in prep.

- in the following assume:
  - $\rightarrow$  LRGs within 0.6 < z < 0.9
  - → CMB lensing with SPT-like noise
  - $\rightarrow$  20 arcmin smoothing scale
  - → 5000 square degrees on the sky

### Data vector:

- $\rightarrow$  pixelizing PDF, 20x20
- $\rightarrow$  use pixels that include 95%
  - of probability

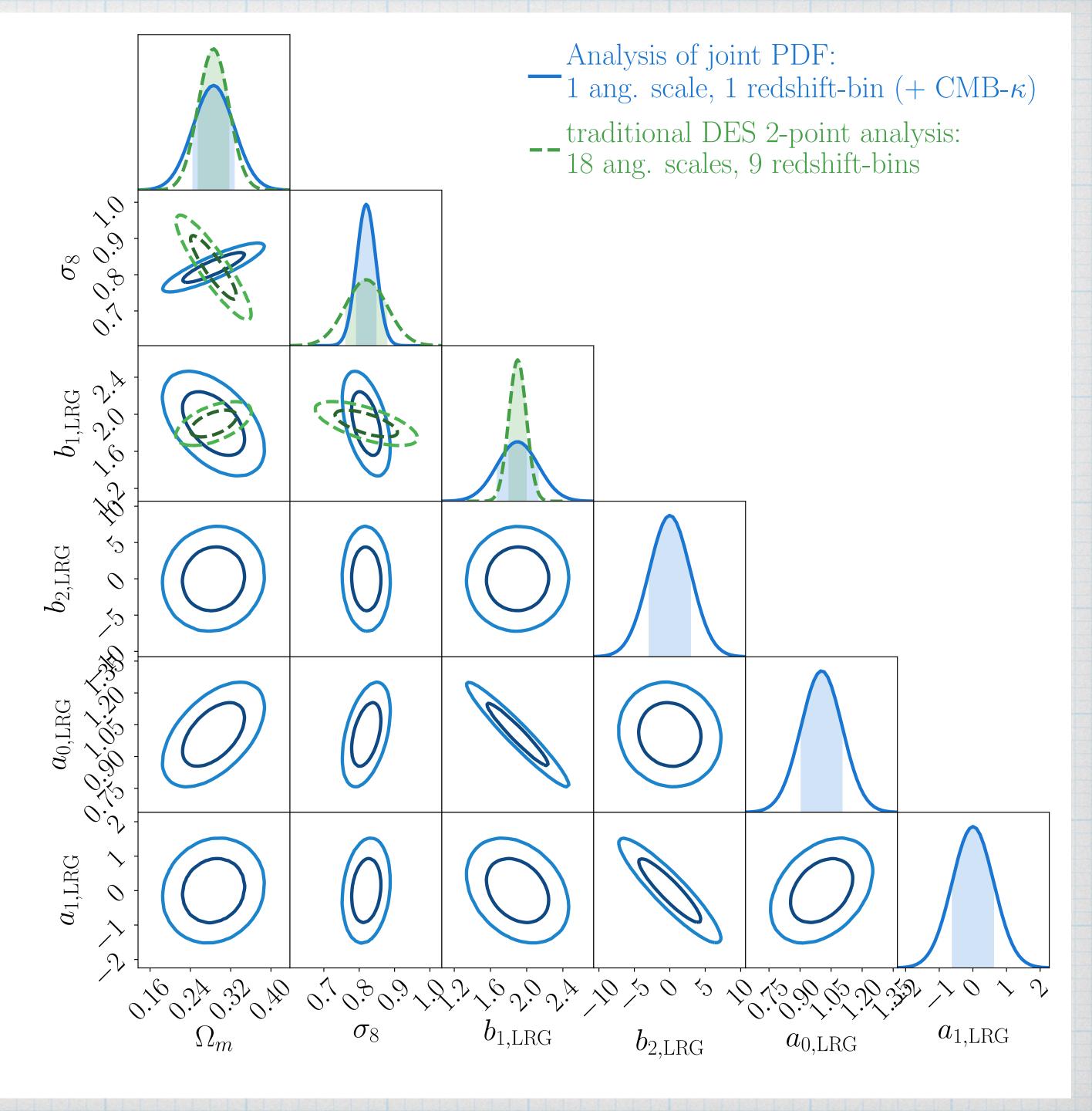
0.152 0.122 0.091 0.061-0.030 KCMB 0.000 -0.030 -0.061 -0.091 -0.122 --0.152

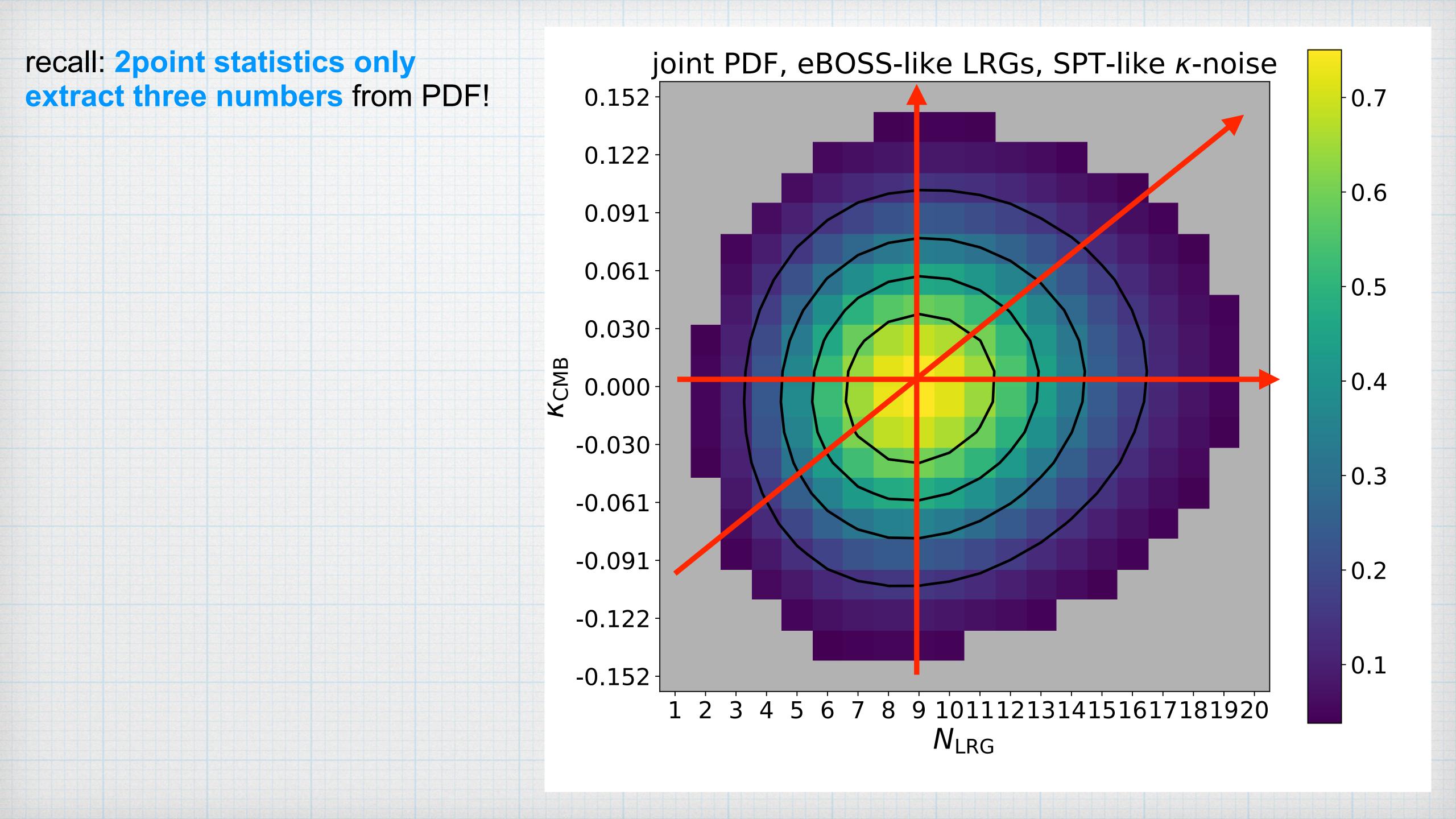


#### PDF can measure

- Omega\_m
- sigma\_8
- linear & quadratic bias
- two parameters of shot-noise

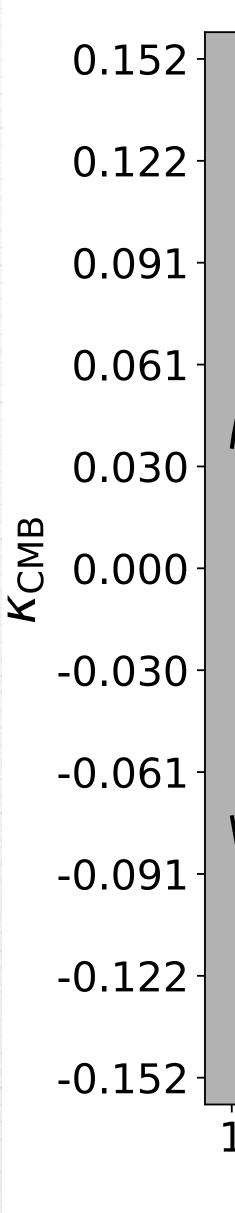
# using only one angular scale & 1 redshift bin!

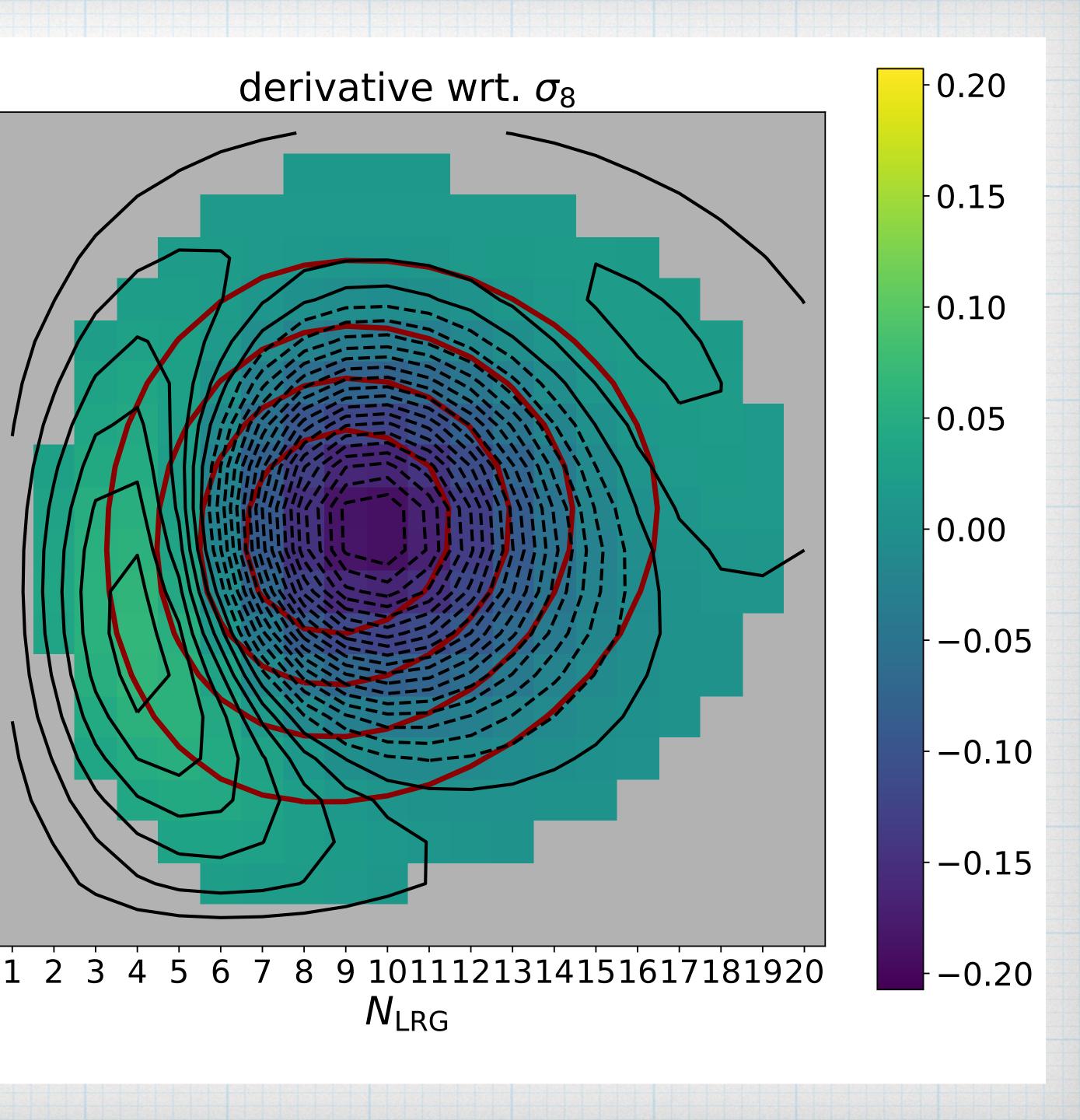




### Red contours: PDF

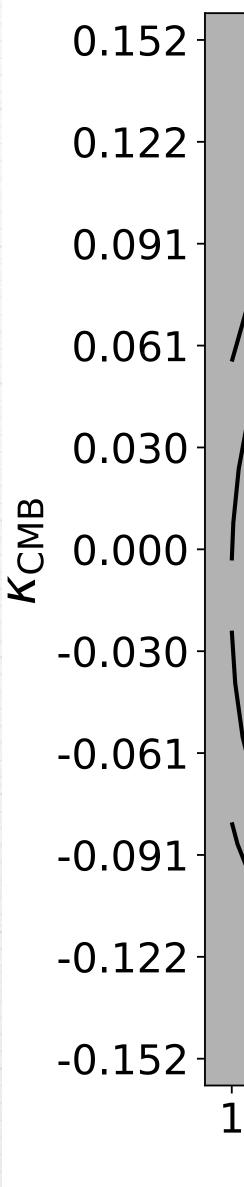
Black contours and colormap: derivative of PDF wrt. different parameters

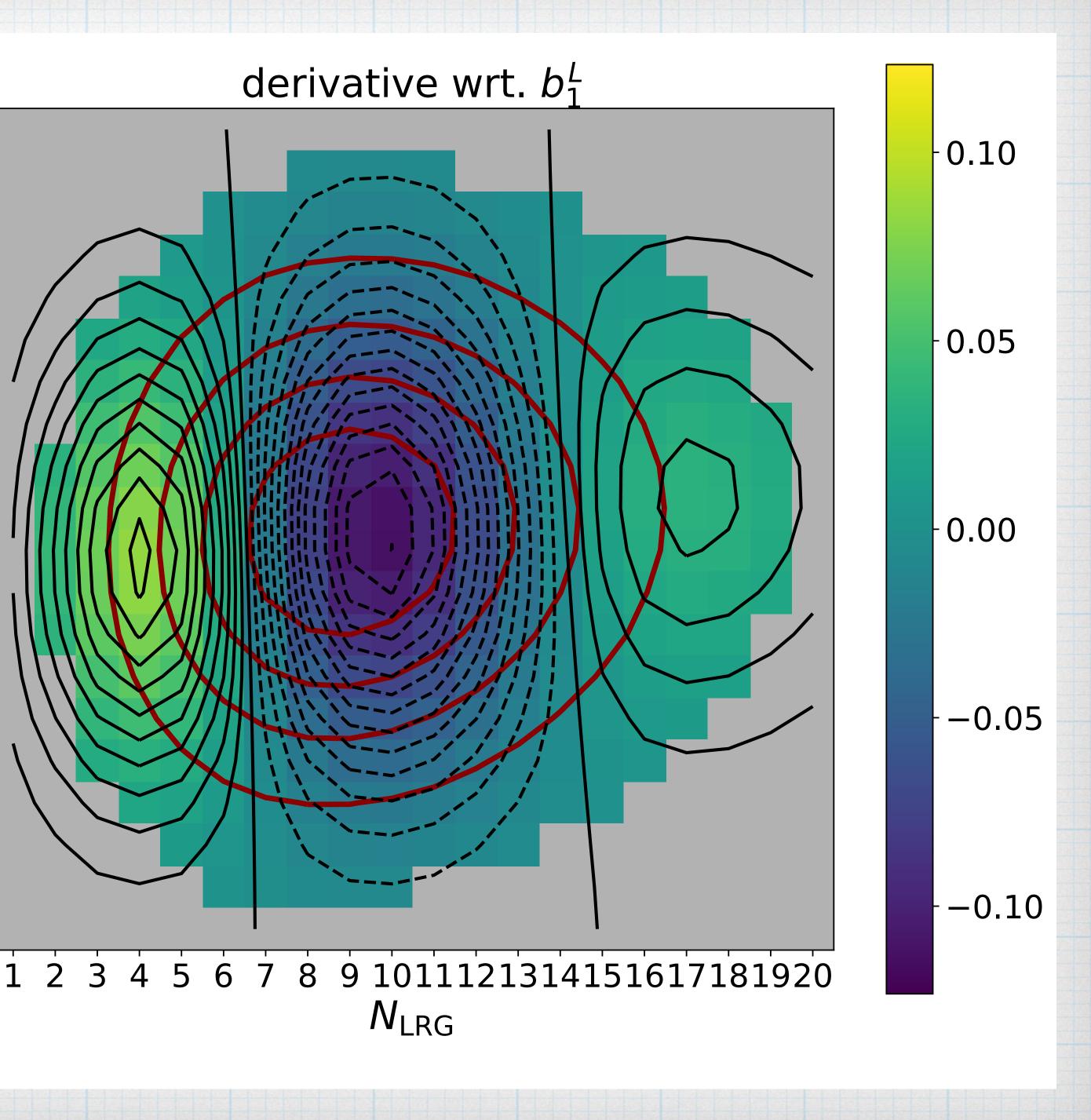




### Red contours: PDF

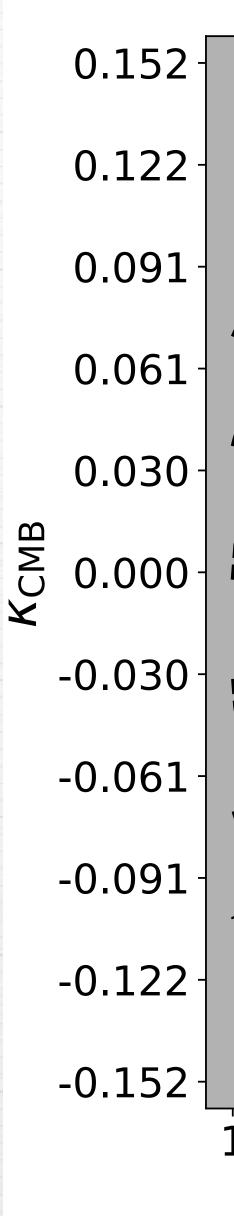
Black contours and colormap: derivative of PDF wrt. different parameters

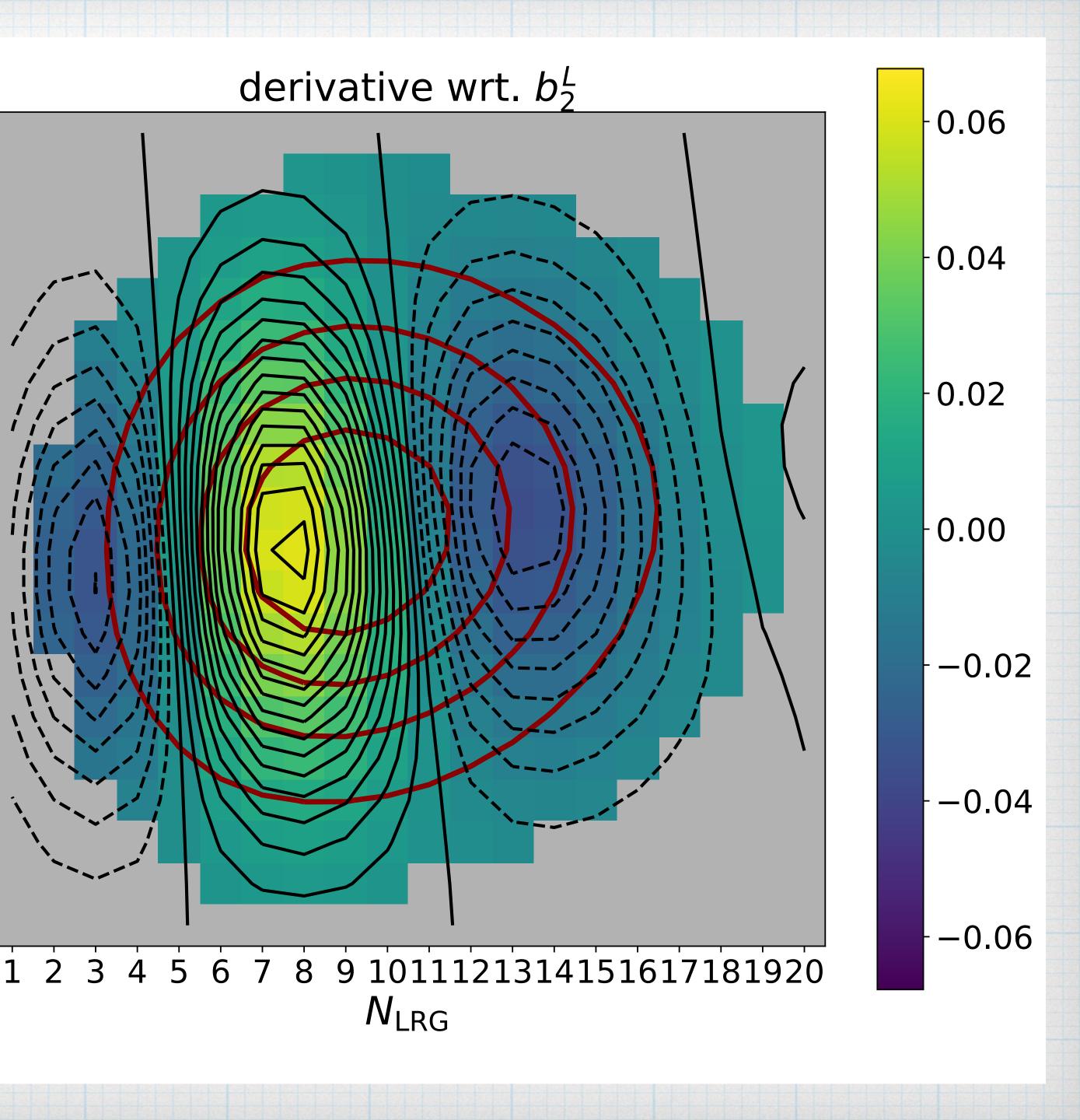


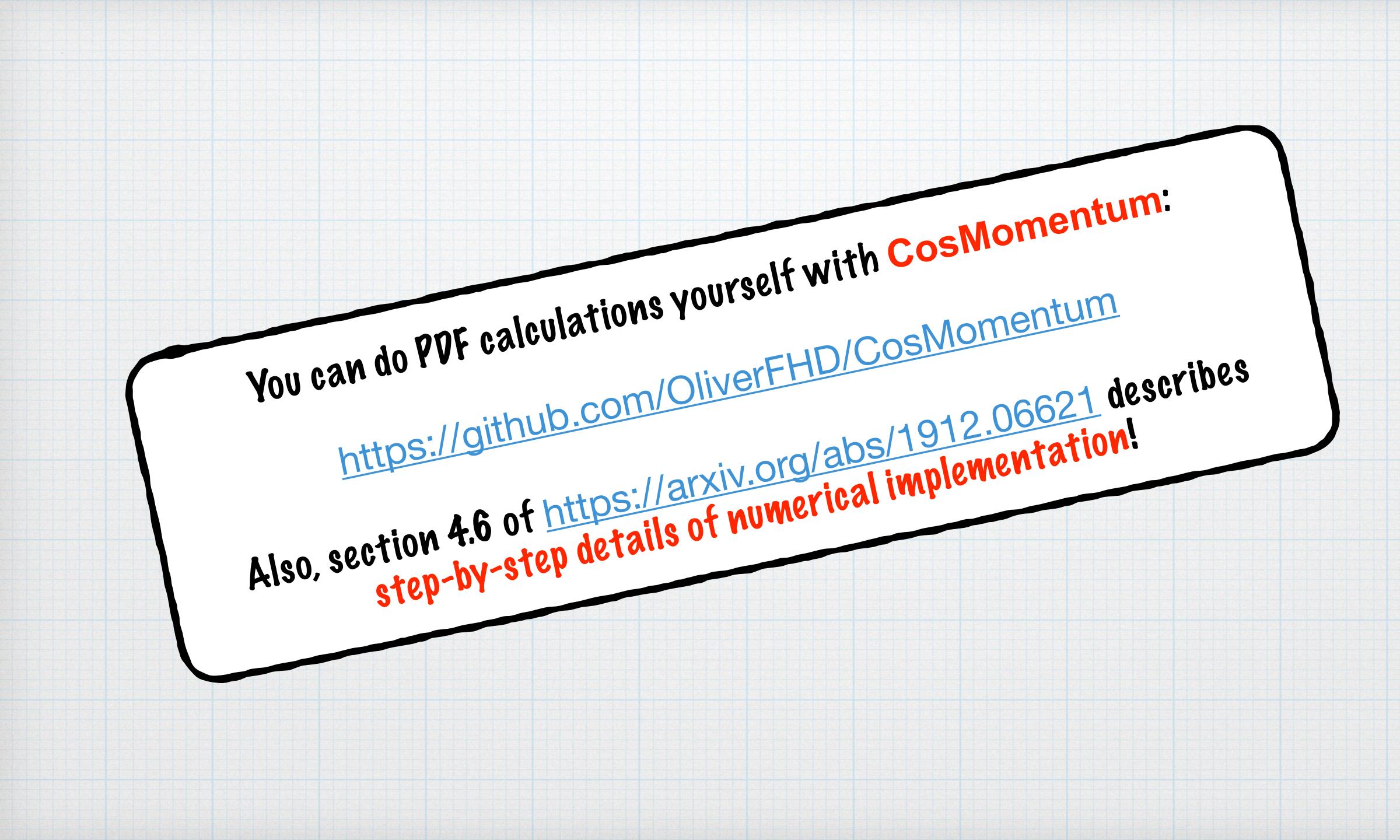


### Red contours: PDF

Black contours and colormap: derivative of PDF wrt. different parameters









### Conclusions & Outlook

PDF technology is catching up with 2-point statistics

 $\rightarrow \mathsf{PNG}$ 

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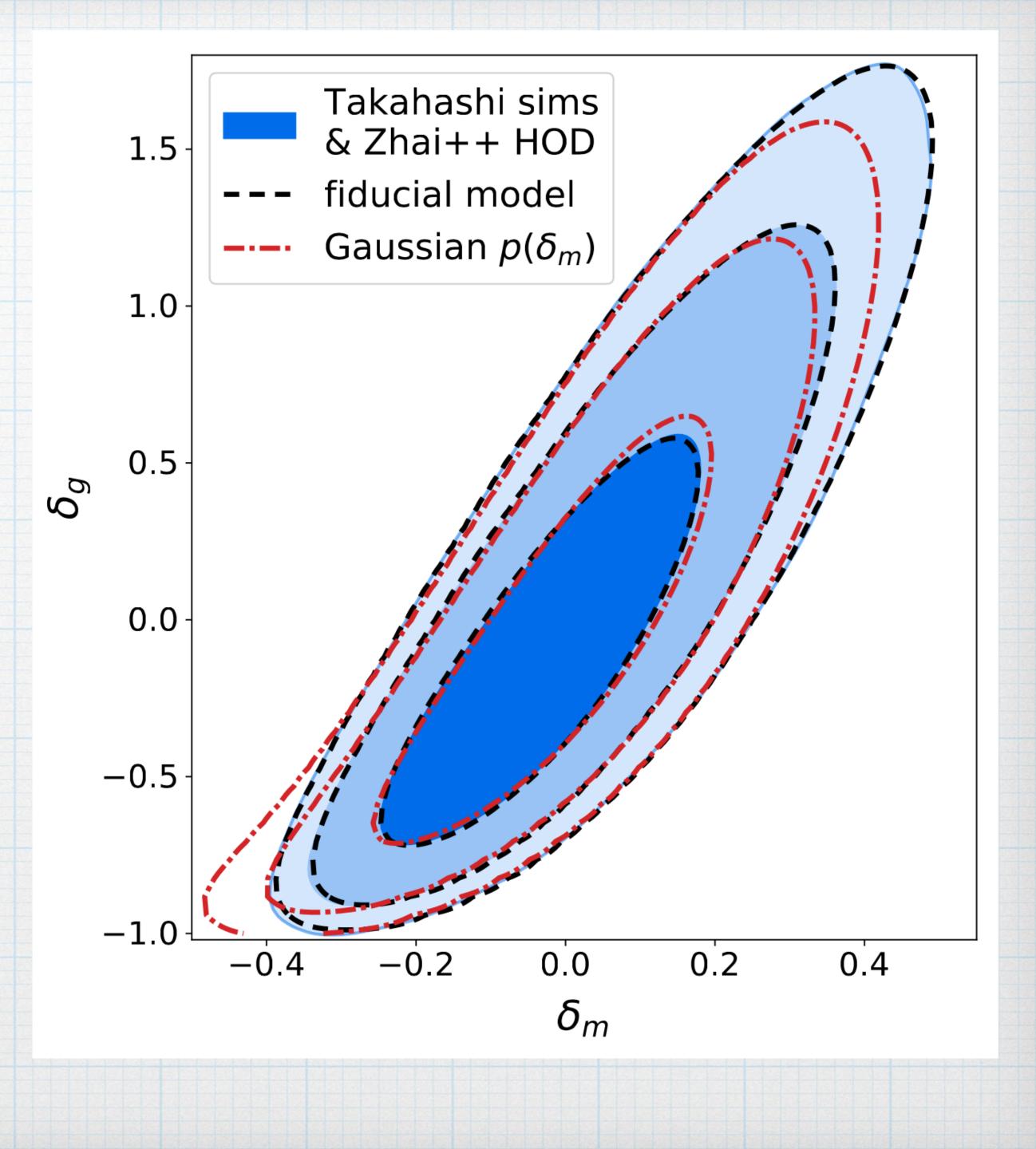
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- $\rightarrow$  Lagrangian bias expansion
- → Neutrino and DE physics

Full shape of PDF very powerful in measuring cosmology → even with 4-parameter bias model!

Potential to break degeneracies of 3x2pt.
What I expect to be most promising:
2-point functions + PDF on one scale

Many of our tools are already publicly available in the **CosMomentum** package: <u>https://github.com/OliverFHD/</u> <u>CosMomentum</u>



## Conclusions & Outlook

- PDF technology is catching up with 2-point statistics
  - $\rightarrow$  PNG

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- → Lagrangian bias expansion
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- Full shape of PDE measu  $\rightarrow ev$ Potent What I

2-point

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