

Next generation cosmological analysis with nested sampling

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8th September 2022



The
Alan Turing
Institute



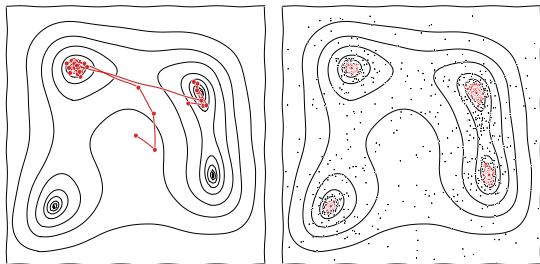
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DiRAC

- ▶ DiRAC 2020 RAC allocation of 30MCPUh
- ▶ Main goal: Planck Legacy Archive equivalent
 - ▶ Parameter estimation \rightarrow Model comparison
 - ▶ MCMC \rightarrow Nested sampling
 - ▶ Planck \rightarrow {Planck, DESY1, BAO, ...}
 - ▶ Pairwise combinations
- ▶ Suite of tools for processing these
 - ▶ unimpeded 1.0
 - ▶ margarine 1.0
 - ▶ anesthetic 2.0
 - ▶ zenodo archive
- ▶ MCMC chains also available.
- ▶ Work in progress, but α -testers requested (email wh260@cam.ac.uk)

DiRAC



The three pillars of Bayesian inference

Parameter estimation

What do the data tell us about the parameters of a model?

e.g. the size or age of a Λ CDM universe

$$P(\theta|D, M) = \frac{P(D|\theta, M)P(\theta|M)}{P(D|M)},$$

$$\mathcal{P} = \frac{\mathcal{L} \times \pi}{\mathcal{Z}},$$

$$\text{Posterior} = \frac{\text{Likelihood} \times \text{Prior}}{\text{Evidence}}.$$

Model comparison

How much does the data support a particular model?
e.g. Λ CDM vs a dynamic dark energy cosmology

$$P(M|D) = \frac{P(D|M)P(M)}{P(D)},$$

$$\frac{\mathcal{Z}_M \Pi_M}{\sum_m \mathcal{Z}_m \Pi_m},$$

$$\text{Posterior} = \frac{\text{Evidence} \times \text{Prior}}{\text{Normalisation}}.$$

Tension quantification

Do different datasets make consistent predictions from the same model?

e.g. CMB vs Type IA supernovae data

$$\mathcal{R} = \frac{\mathcal{Z}_{AB}}{\mathcal{Z}_A \mathcal{Z}_B},$$

$$\begin{aligned} \log \mathcal{S} = & \langle \log \mathcal{L}_{AB} \rangle_{\mathcal{P}_{AB}} \\ & - \langle \log \mathcal{L}_A \rangle_{\mathcal{P}_A} \\ & - \langle \log \mathcal{L}_B \rangle_{\mathcal{P}_B} \end{aligned}$$

Occam's Razor [2102.11511]

- ▶ Bayesian inference quantifies Occam's Razor:
 - ▶ *"Entities are not to be multiplied without necessity"* — William of Occam
 - ▶ *"Everything should be kept as simple as possible, but not simpler"* — "Albert Einstein"
- ▶ Properties of the evidence: rearrange Bayes' theorem for parameter estimation

$$\mathcal{P}(\theta) = \frac{\mathcal{L}(\theta)\pi(\theta)}{\mathcal{Z}} \Rightarrow \log \mathcal{Z} = \log \mathcal{L}(\theta) - \log \frac{\mathcal{P}(\theta)}{\pi(\theta)}$$

- ▶ Evidence is composed of a "goodness of fit" term and "Occam Penalty"
- ▶ RHS true for all θ . Take max likelihood value θ_* :
- ▶ Be more Bayesian and take posterior average to get the "Occam's razor equation"

$$\log \mathcal{Z} = -\chi_{\min}^2 - \text{Mackay penalty}$$

$$\log \mathcal{Z} = \langle \log \mathcal{L} \rangle_{\mathcal{P}} - \mathcal{D}_{\text{KL}}$$

- ▶ Natural regularisation which penalises models with too many parameters.

Kullback Liebler divergence

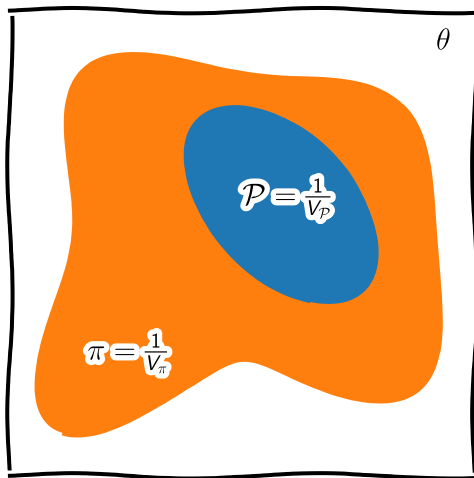
- ▶ The KL divergence between **prior** π and **posterior** \mathcal{P} is defined as:

$$\mathcal{D}_{\text{KL}} = \left\langle \log \frac{\mathcal{P}}{\pi} \right\rangle_{\mathcal{P}} = \int \mathcal{P}(\theta) \log \frac{\mathcal{P}(\theta)}{\pi(\theta)} d\theta.$$

- ▶ Whilst not a distance, $\mathcal{D} = 0$ when $\mathcal{P} = \pi$.
- ▶ Occurs in the context of machine learning as an objective function for training functions.
- ▶ In Bayesian inference it can be understood as a log-ratio of “volumes”:

$$\mathcal{D}_{\text{KL}} \approx \log \frac{V_{\pi}}{V_{\mathcal{P}}}.$$

(this is exact for top-hat distributions).

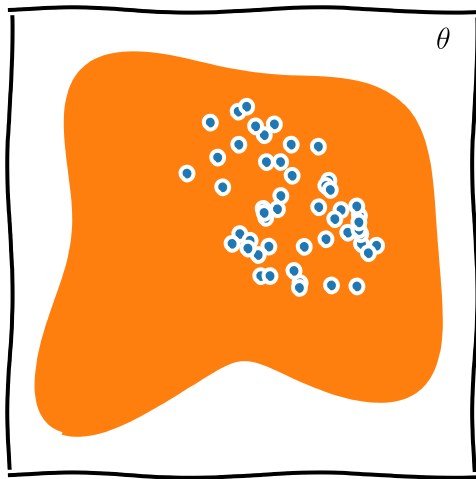


Why do sampling?

- ▶ The cornerstone of numerical Bayesian inference is working with **samples**.
- ▶ Generate a set of representative parameters drawn in proportion to the posterior $\theta \sim \mathcal{P}$.
- ▶ The magic of marginalisation \Rightarrow perform usual analysis on each sample in turn.
- ▶ The golden rule is **stay in samples** until the last moment before computing summary statistics/triangle plots because

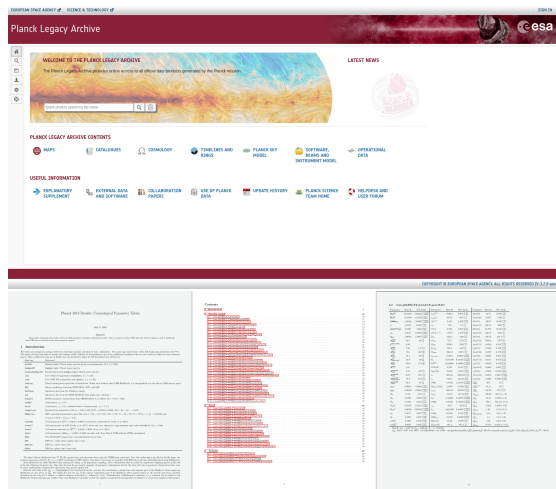
$$f(\langle X \rangle) \neq \langle f(X) \rangle$$

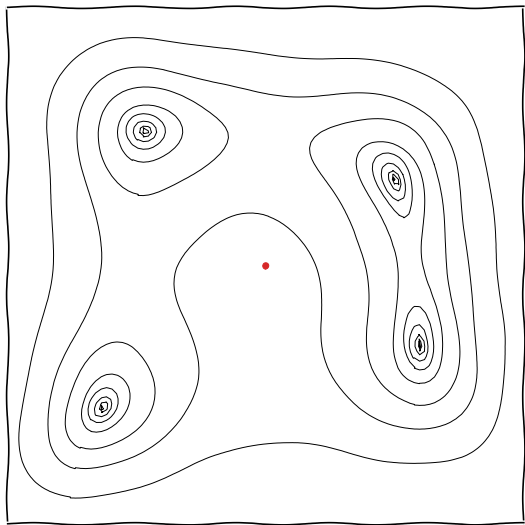
- ▶ Generally need $\sim \mathcal{O}(12)$ independent samples to compute a value and error bar.

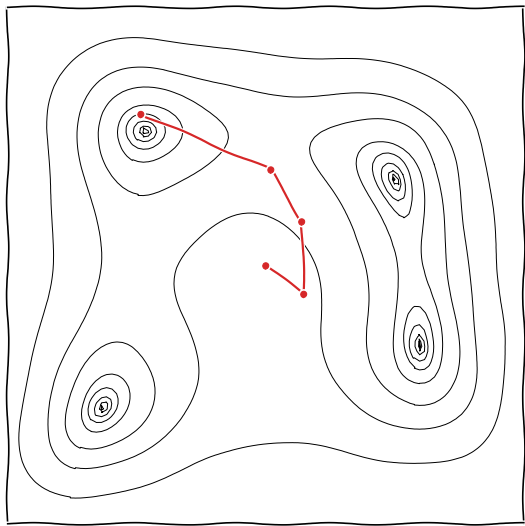


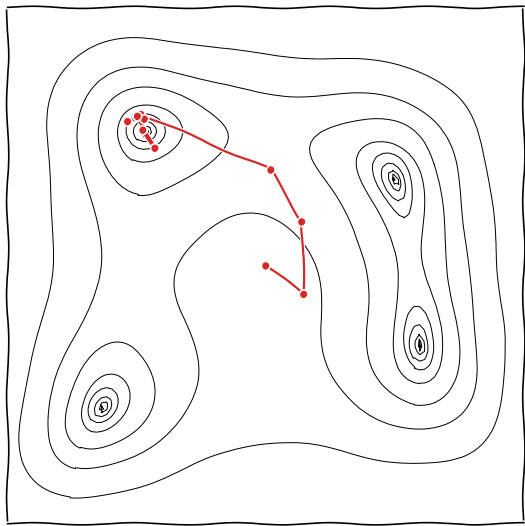
The Planck legacy archive

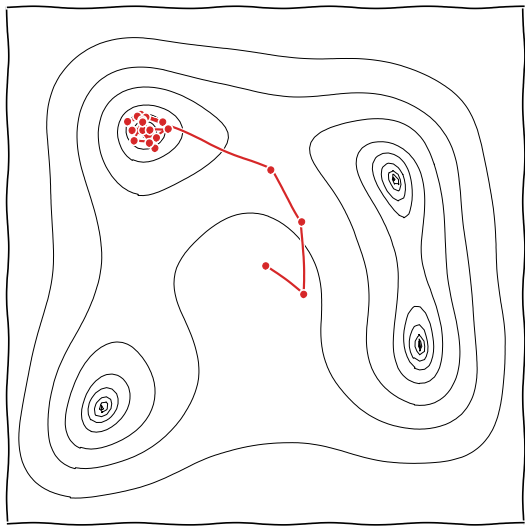
- ▶ *Planck* collaboration science products
- ▶ Distributed cosmology inference results as MCMC chains
- ▶ Across a grid of:
 - ▶ subsets/combinations of *Planck* data
 - ▶ TT, lowl, lowE, lensing
 - ▶ Λ CDM extensions
 - ▶ base, mnu, nrun, ω_{de} , r
- ▶ importance sampling across some other likelihoods (BAO, JLA, ...)
- ▶ Cannot compute evidences in high dimensions from MCMC chains
 - ▶ Only parameter estimation
 - ▶ no model comparison

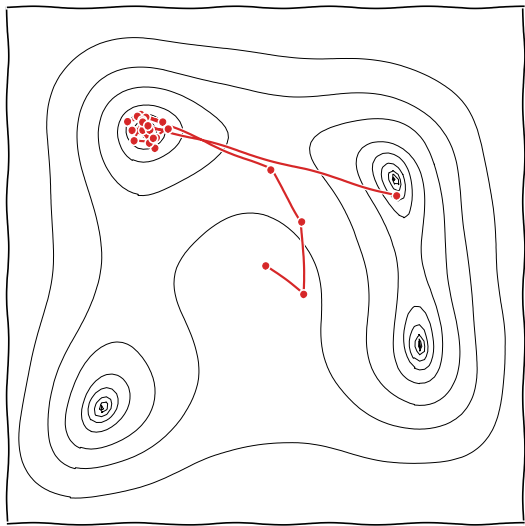


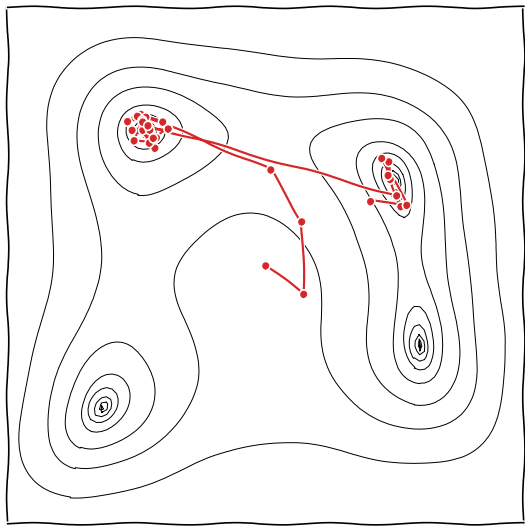




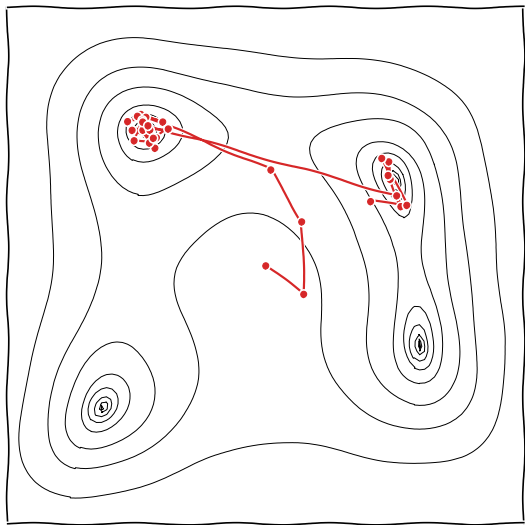




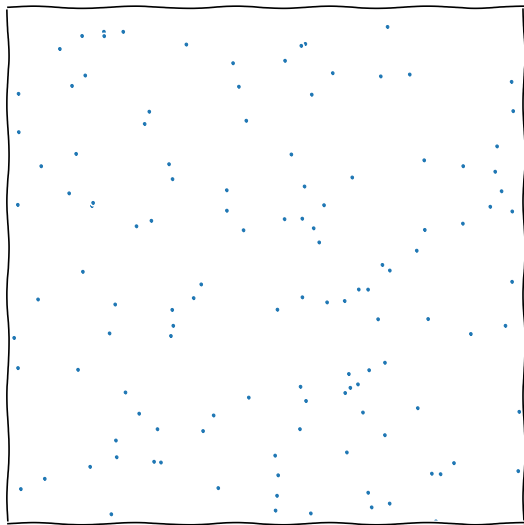




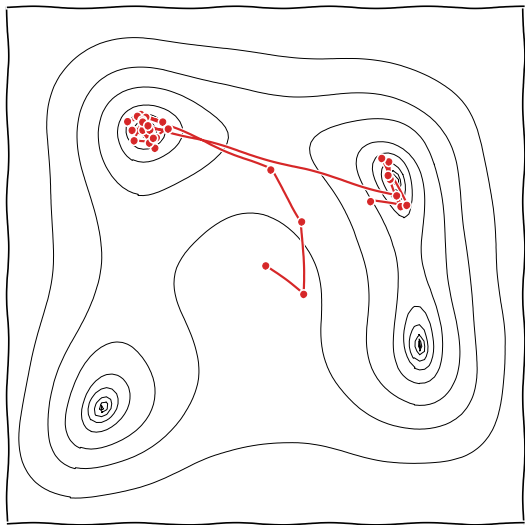
MCMC



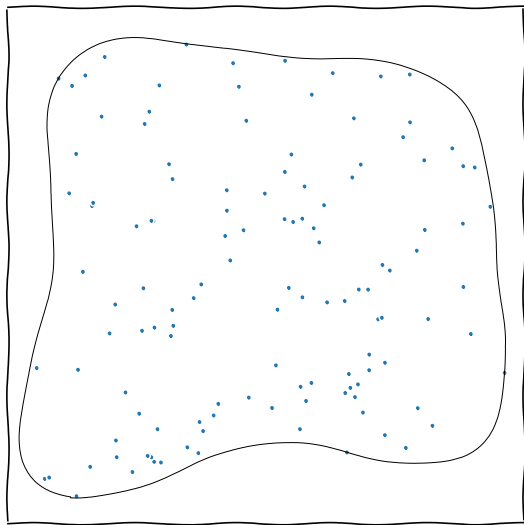
Nested sampling

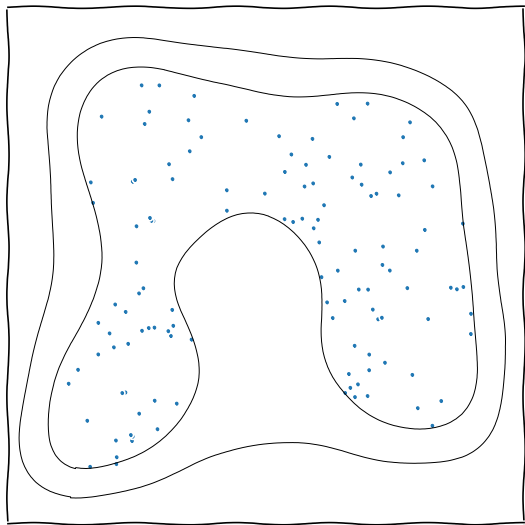
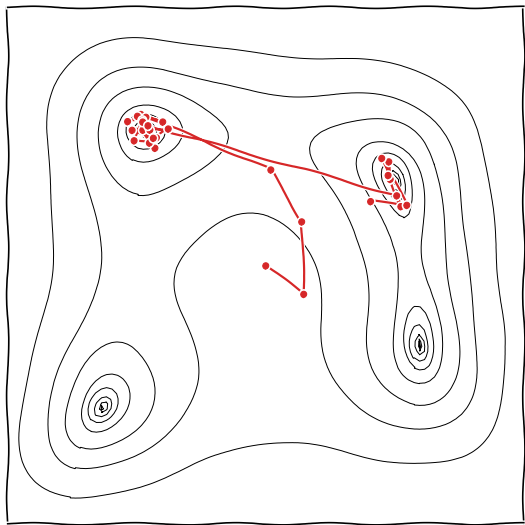


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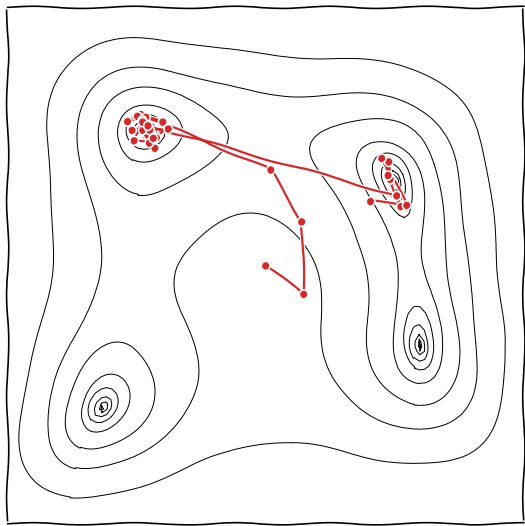


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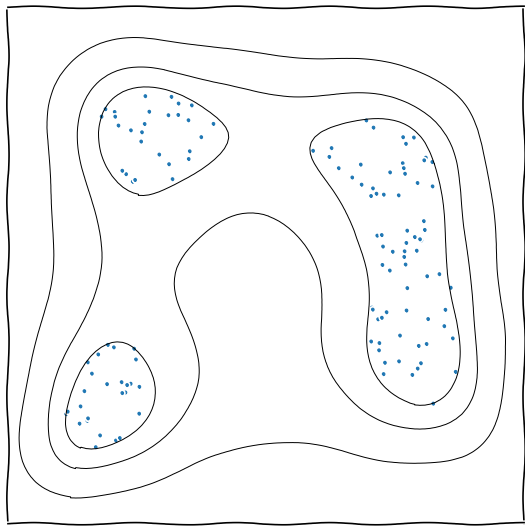




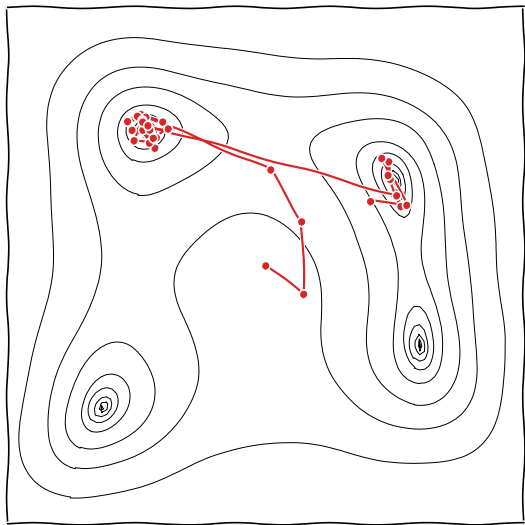
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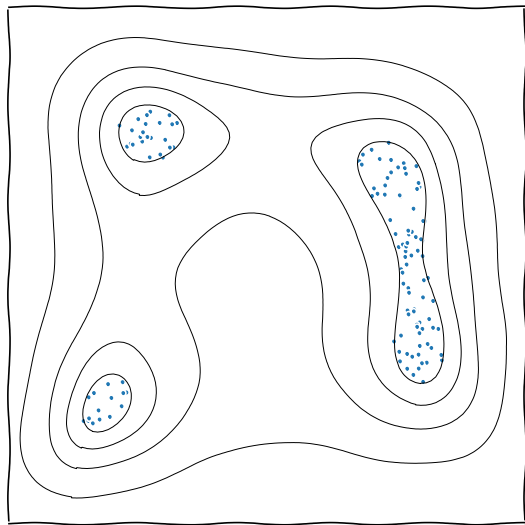
Nested sampling



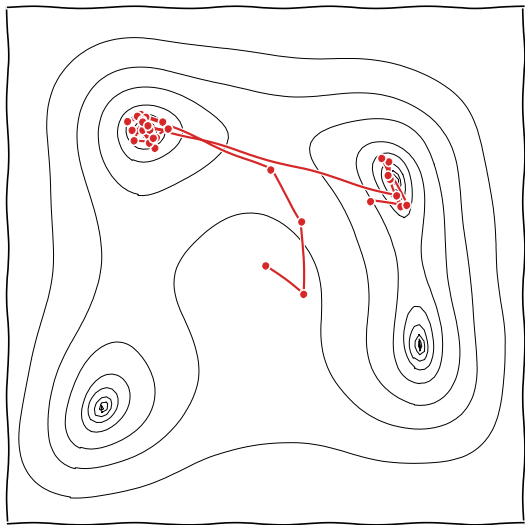
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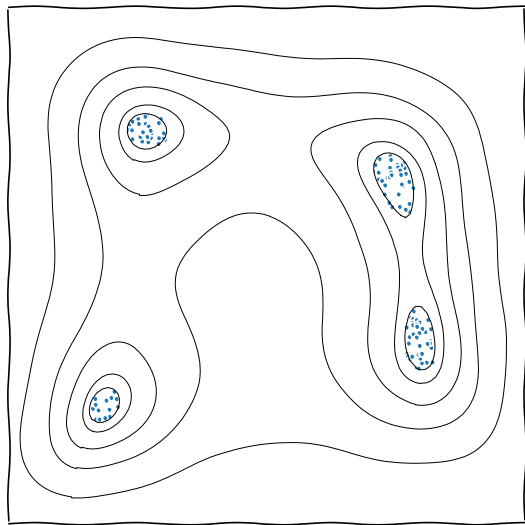
Nested sampling



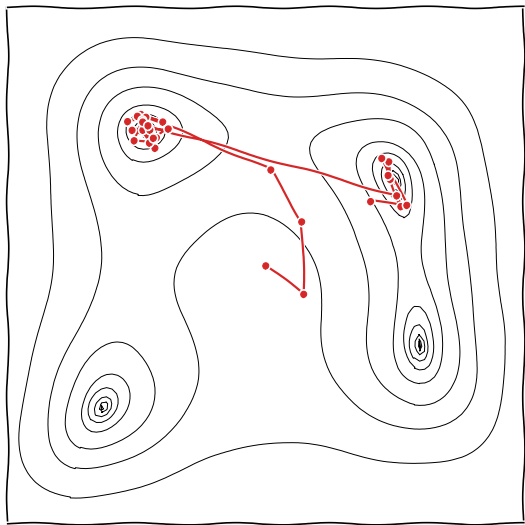
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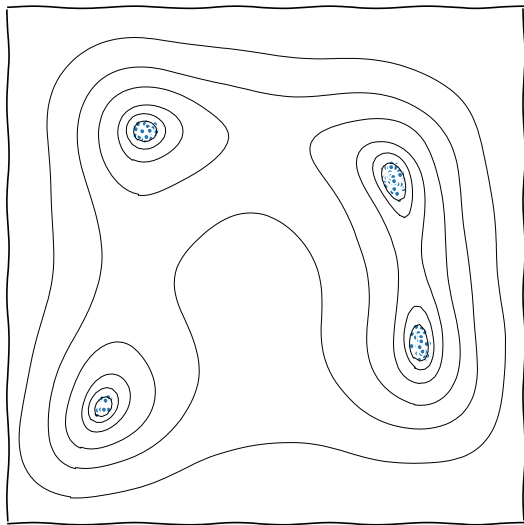
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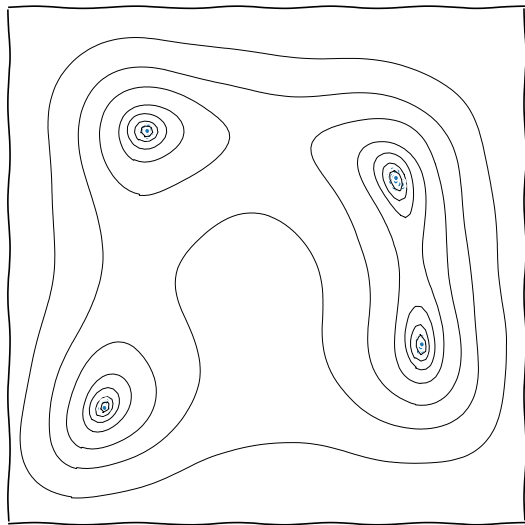
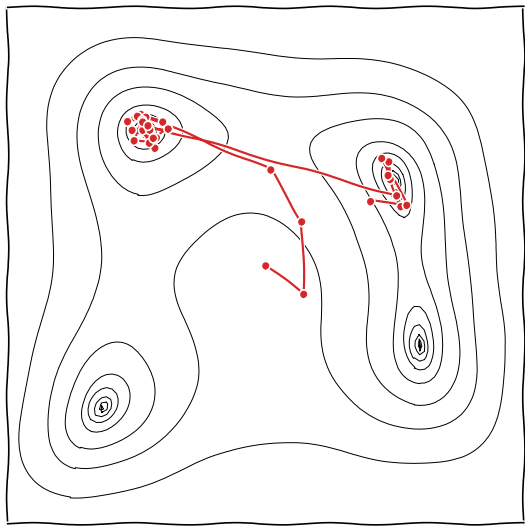


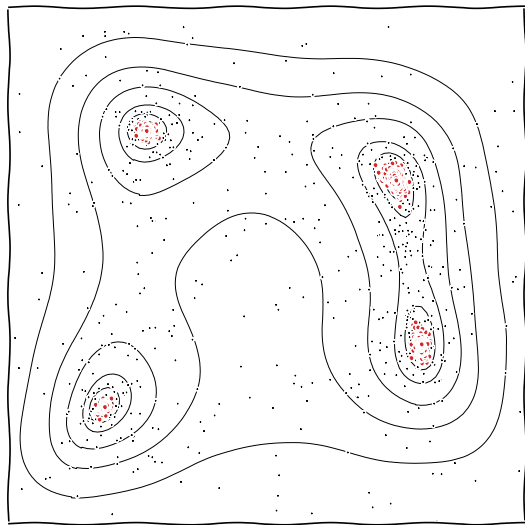
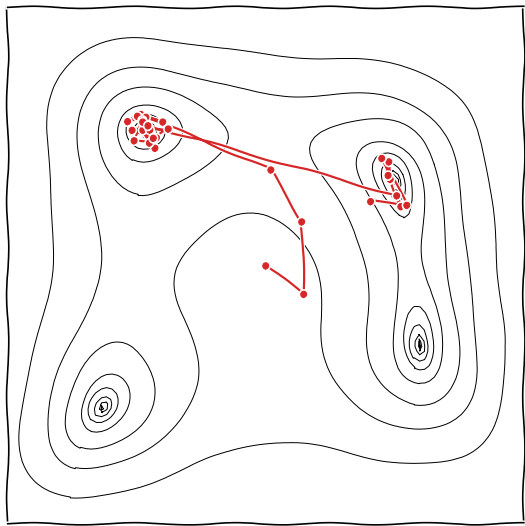
MCMC



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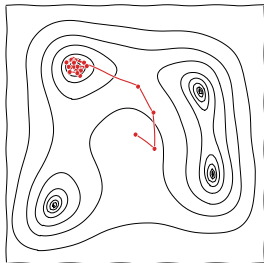






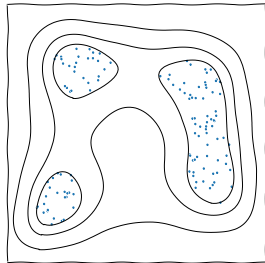
MCMC

- ▶ Single “walker”
- ▶ Explores posterior
- ▶ Fast, if proposal matrix is tuned
- ▶ Parameter estimation, suspiciousness calculation
- ▶ Channel capacity optimised for generating posterior samples



Nested sampling

- ▶ Ensemble of “live points”
- ▶ Scans from prior to peak of likelihood
- ▶ Slower, no tuning required
- ▶ Parameter estimation, model comparison, tension quantification
- ▶ Channel capacity optimised for computing partition function

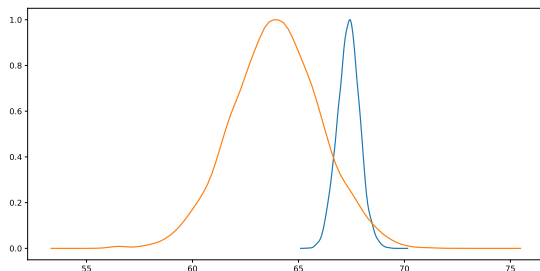


The grid (so far)

- ▶ Models: $[\Lambda\text{CDM}, \Omega_K, \nu, r, w, w(a)]$
- ▶ Data: [plik, camspec, DESY1, bicep+keck, BAO(DR16), pantheon]
- ▶ Pairwise combinations of datasets
- ▶ Breakdown of Planck & BAO data
- ▶ Samplers: [Metropolis Hastings MCMC, Nested Sampling]
- ▶ These exhaust what is currently available by default in cobaya
- ▶ Wide priors to allow for importance readjustment as desired
- ▶ roughly halfway through computational allocation.
- ▶ Feedback desirable as to what extensions to the grid would be of community interest (email wh260@cam.ac.uk) (Pantheon+, SH_0 ES, NPIPE, DESY3, ...).
- ▶ Further checking needed before first release by end of this year.

- ▶ Python tool for seamlessly downloading, uploading and cacheing of chains
- ▶ Data stored on zenodo
- ▶ hdf5 storage for fast & reliable storage
- ▶ anesthetic compatible for processing of chains [1905.04768]
- ▶ α -testers wanted! (email wh260@cam.ac.uk)
- ▶ End goal – community library which everyone contributes to so expensive inference products are reusable and reused.

```
from unimpeded import Unimpeded
store = Unimpeded(cache='data.hdf5')
samps = store('planck')
samps.H0.plot.kde_1d()
samps = store('planck', model='klcdm')
samps.H0.plot.kde_1d()
```





Harry Bevens
[2205.12841]
[2207.11457]

- ▶ Can use machine learning + grid to dramatically speed up inference
- ▶ Emulate the marginal posterior and prior with masked autoregressive flows (margarine)
- ▶ Use nested sampling evidences to compute nuisance marginalised likelihood $\mathcal{L}(\theta) = \mathcal{P}(\theta)\mathcal{Z}/\pi(\theta)$

- ▶ Library of pre-trained bijectors to be used as priors/emulators/nuisance marginalised likelihoods
- ▶ e.g. easy to apply a *Planck*/DES/HERA/JWST prior or likelihood to your existing MCMC chains without needing to install the whole cosmology machinery.



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margarine: machine learning-enhanced Bayesian inference

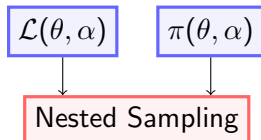


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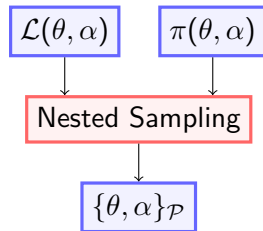


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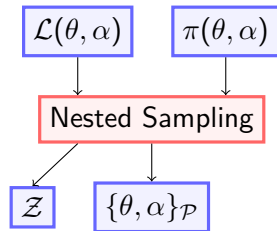


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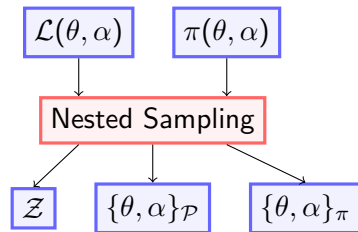
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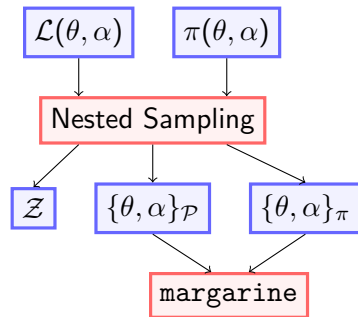
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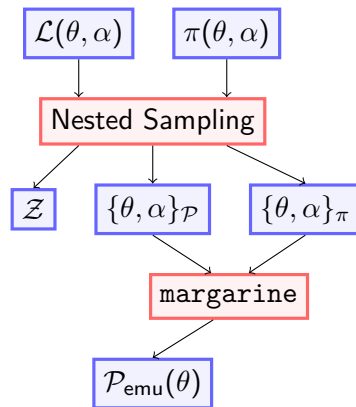


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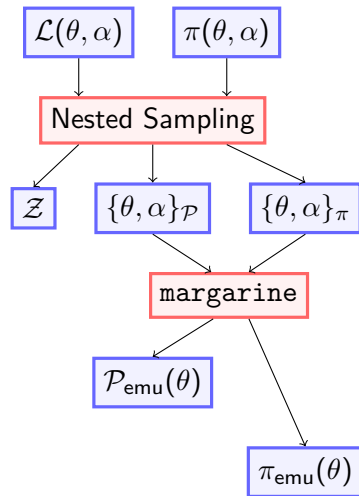


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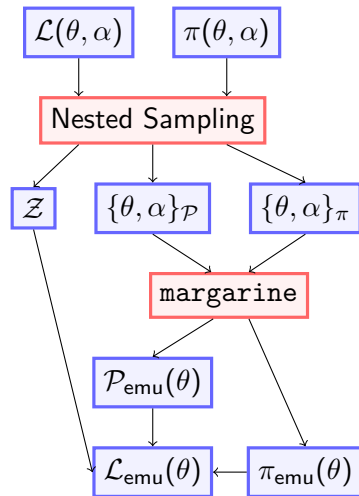


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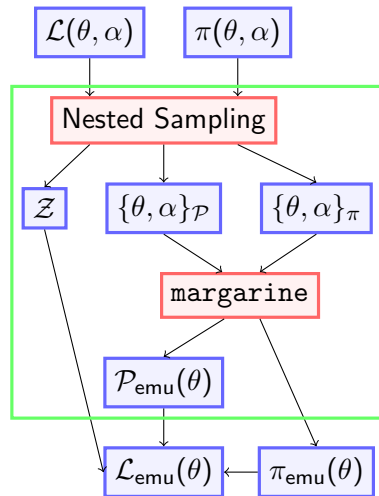


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Conclusions

- ▶ DiRAC RAC allocation for building a legacy grid of
 - ▶ MCMC & Nested sampling chains
 - ▶ gridded over (pairwise) up-to-date datasets
 - ▶ gridded over extensions to Λ CDM
 - ▶ Bijectors & emulators for fast re-use
 - ▶ Importance sampling toolkit via `anesthetic` for (re)processing
 - ▶ Long-term goal: community repository of chains to share model comparison compute resource
- ▶ Looking for:
 - ▶ α -testers for unimpeded
 - ▶ Suggestions for more datasets (and their incorporation into `cobaya`)