



EAPS



MIT KAVLI INSTITUTE

LOW ALBEDO SURFACES OF LAVA WORLDS

ZAHRA ESSACK

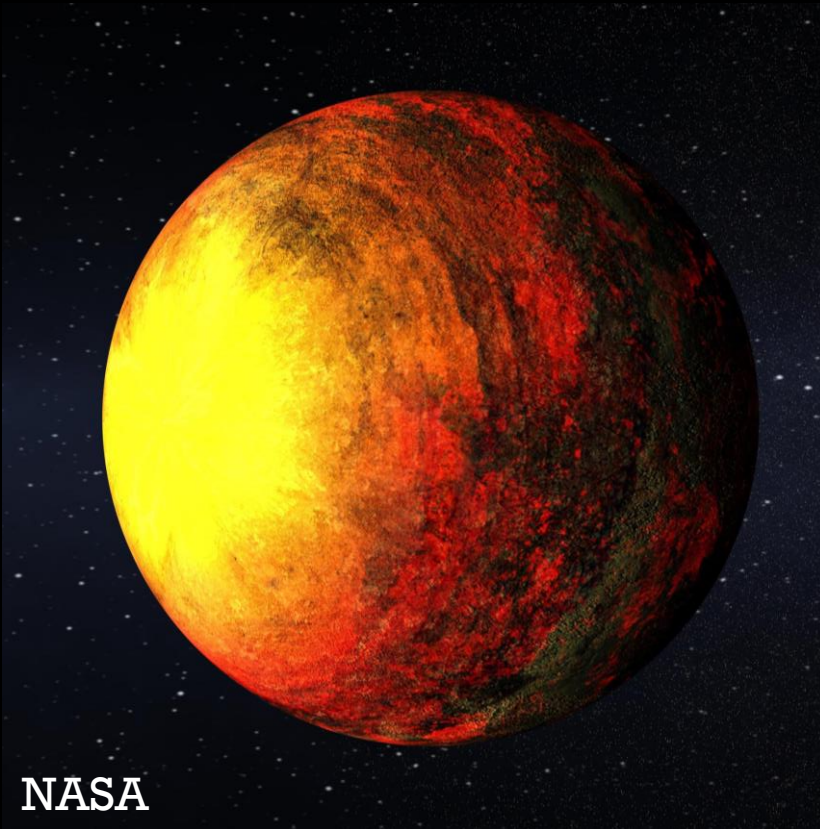
SARA SEAGER, MIHKEL PAJUSALU

ROCKY WORLDS, CAMBRIDGE UK

8 JANUARY 2020

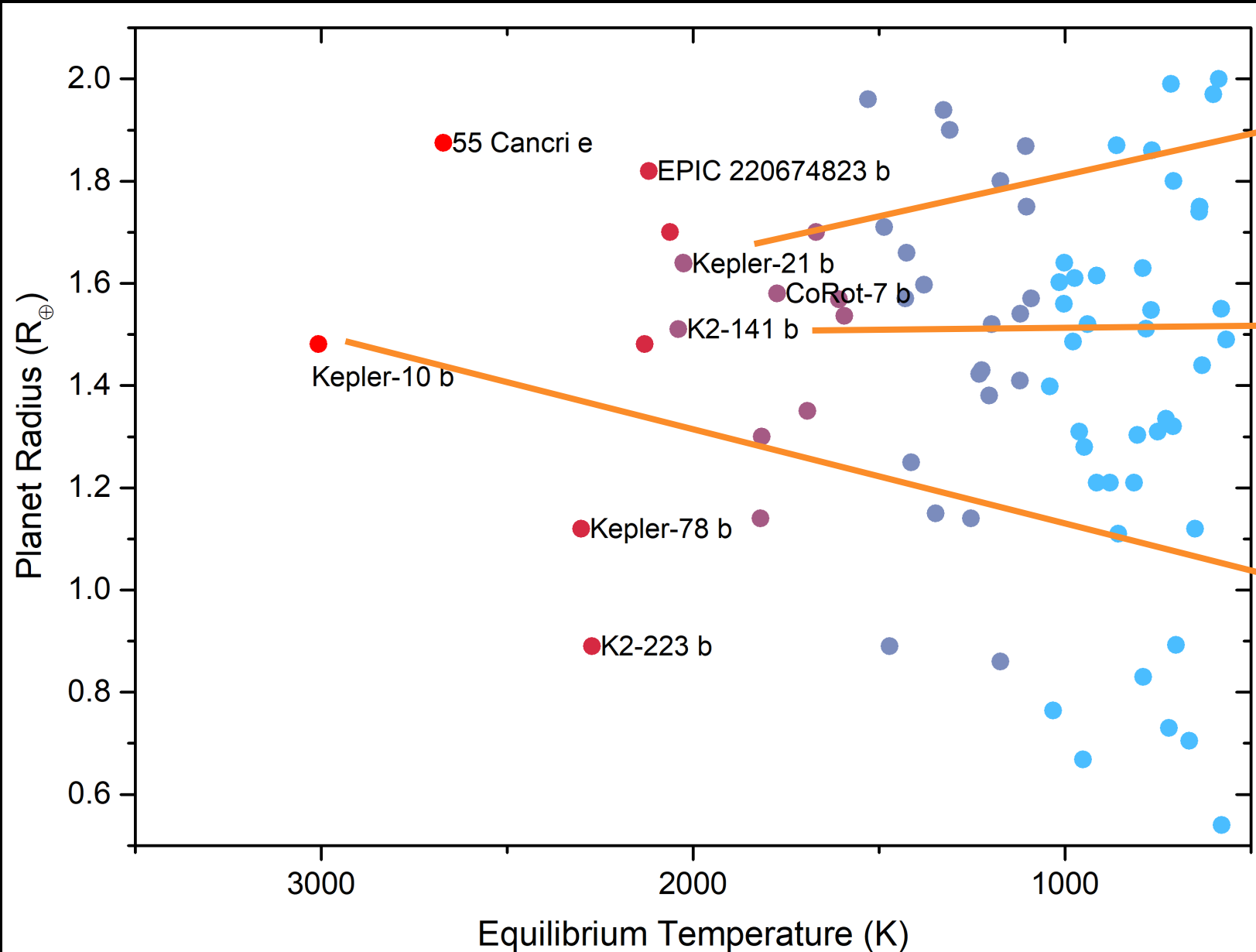
HOT SUPER EARTHS LAVA-OCEAN EXOPLANETS

What causes the high geometric albedos on some hot super Earths?



- $R_{\text{planet}} < 1.6 R_{\text{earth}}$
- Tidally locked
- Low pressure atmospheres (< 0.1 bar)
- Substellar temperature > 850 K
- Surface lava oceans due to intense stellar irradiation

LAVA-OCEAN EXOPLANET CANDIDATES



$$0.4 < A_g < 0.5$$

$$0.2 < A_g < 0.4$$

$$0.4 < A_g < 0.5$$

Demory (2014)

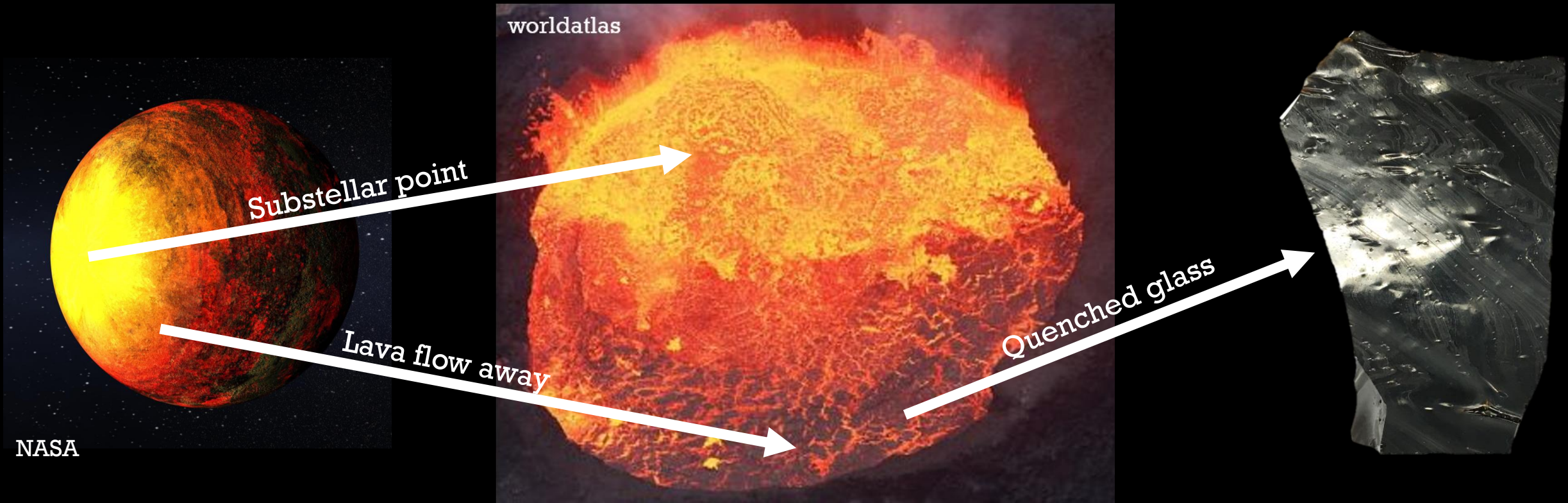
Malavolta et al. (2018)

$A_g = \text{Geometric Albedo}$

SURFACES AS A SOURCE OF HIGH ALBEDOS



A (SIMPLE) THEORETICAL SURFACE OF A LAVA WORLD



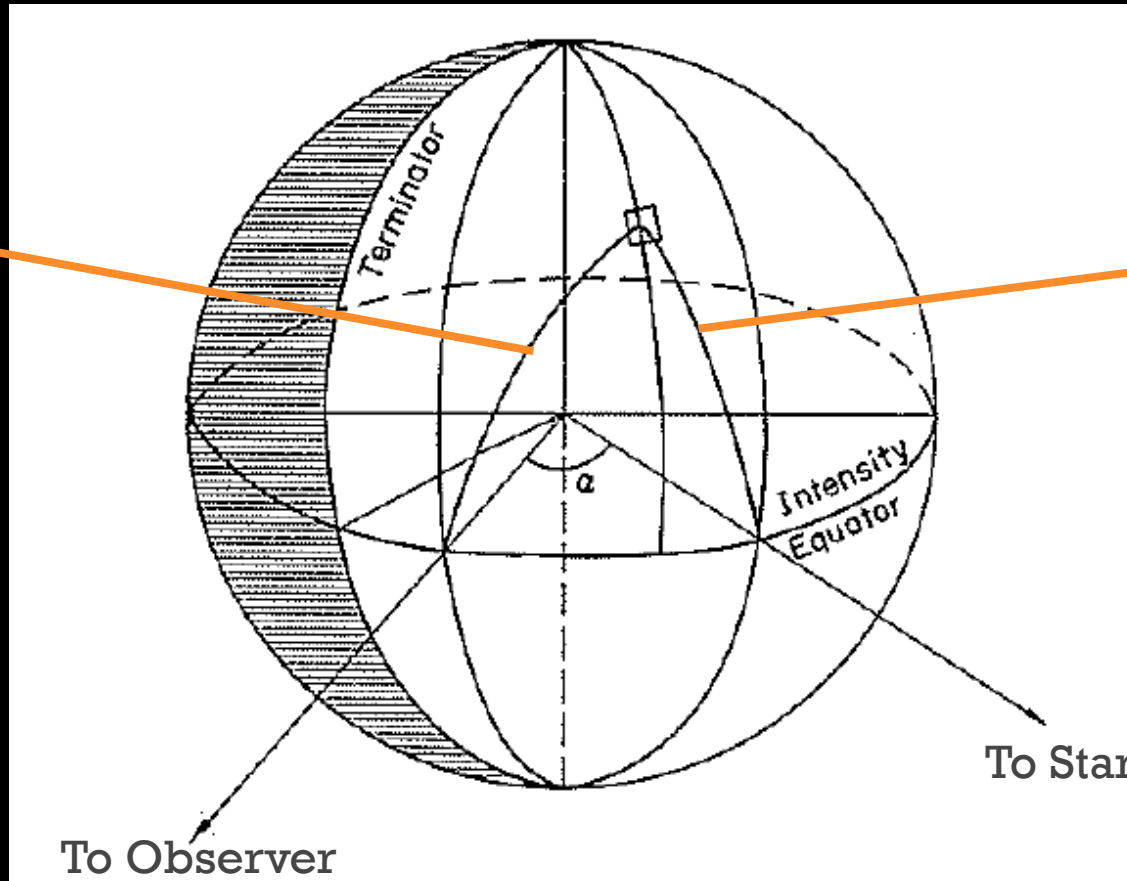
GEOMETRIC ALBEDO OF A PLANET

Reflected angle: η

$$\cos(\text{latitude}) \cos(\text{longitude})$$

Incidence angle ζ

$$\cos(\text{latitude}) \cos(\text{longitude} - \alpha)$$

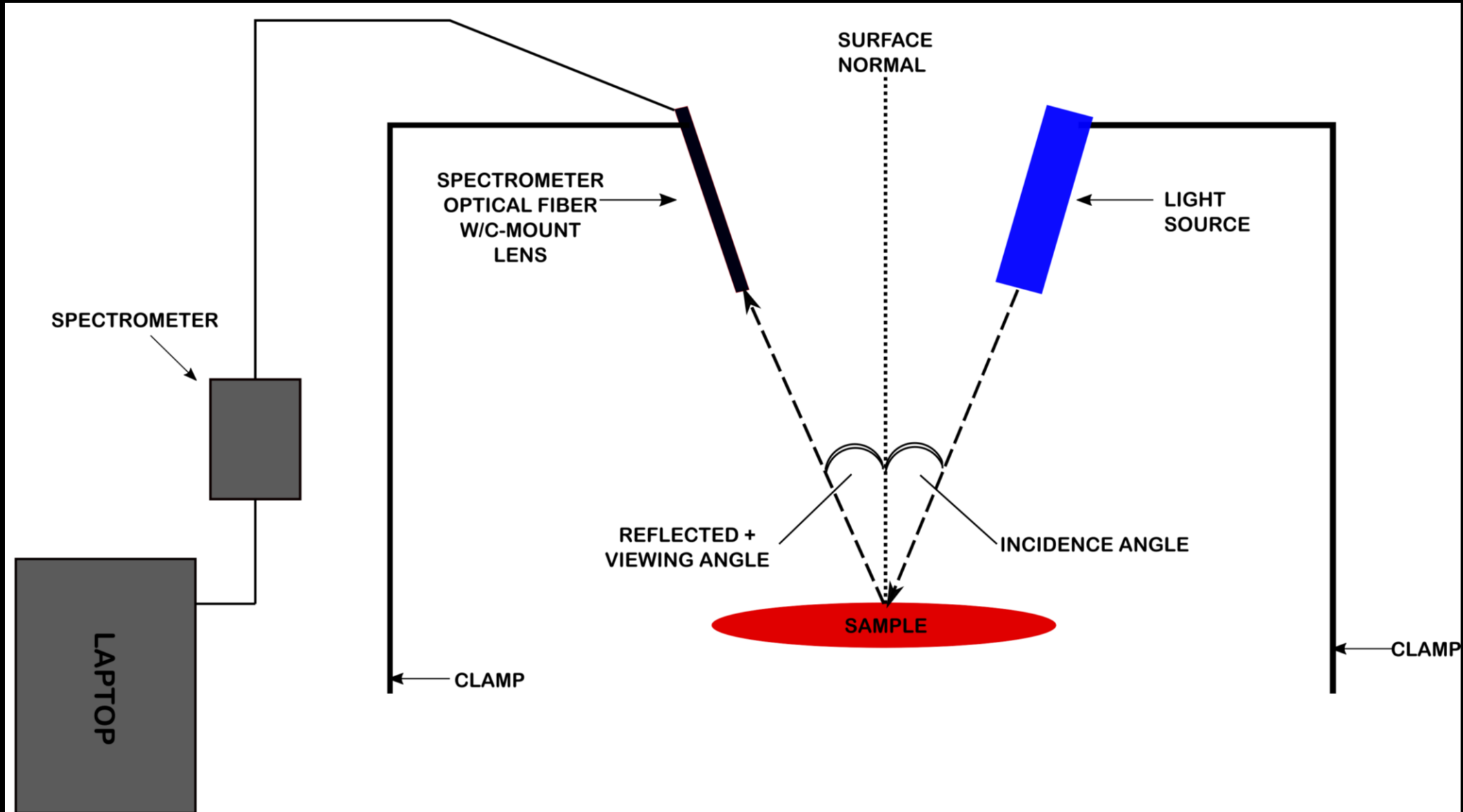


Reflection coefficient

$$\varrho(\eta, \zeta, \varphi)$$

$$A_g = 2 \int_0^1 \varrho(\eta, \eta, \pi) \eta^2 d\eta$$

MEASURING REFLECTION FROM QUENCHED GLASSES

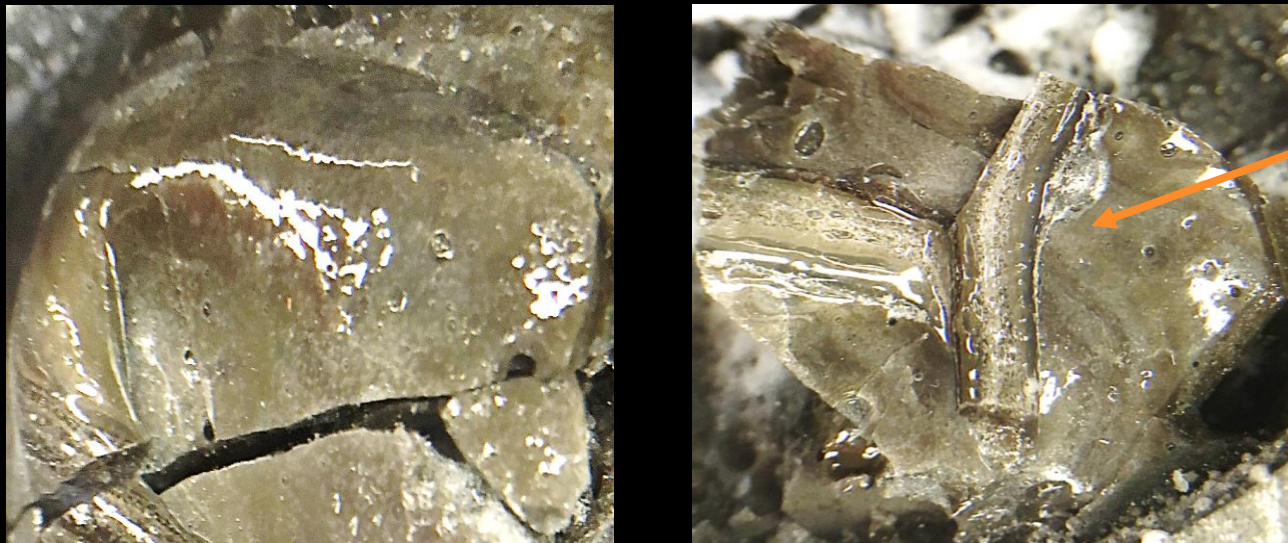


BASALT AND FELDSPAR QUENCHED GLASSES

Basalt



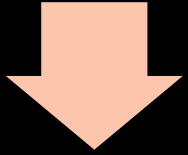
Feldspar



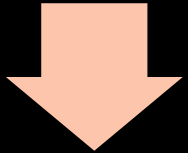
BASALT: REFLECTANCE VS. η (ROUGH GLASS)

Essack et al. (in review)

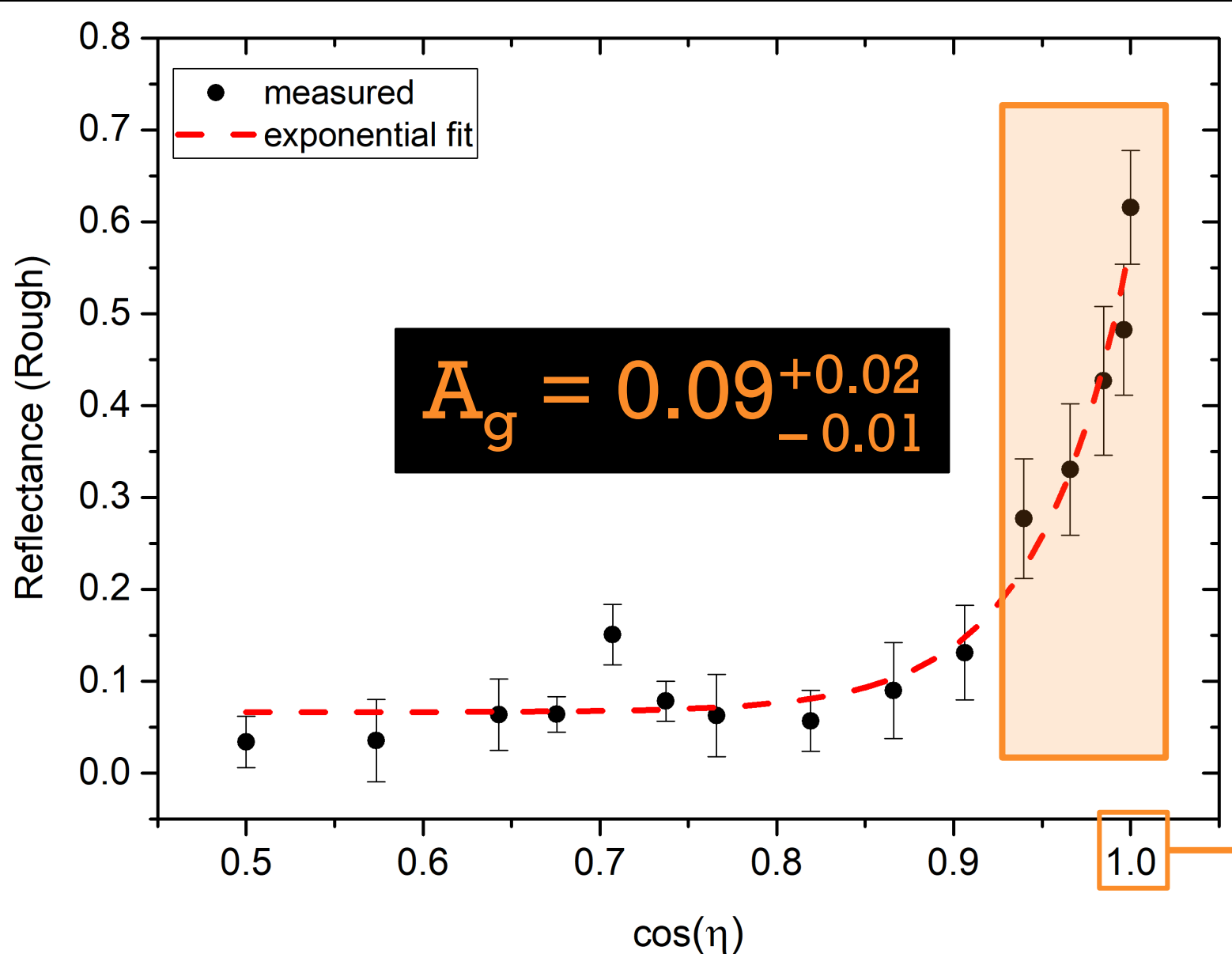
Lab measurements
of reflection from
quenched glass



Fit data from lab
measurements to get
**reflection
coefficient
function: $\rho(\eta, \eta, \pi)$**



Integrate reflection
coefficient function
over all latitudes and
longitudes on the
planet dayside
hemisphere to get
albedo: A_g



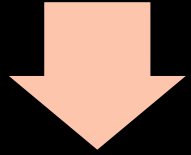
Incidence Angle =
Reflected Angle η

Substellar
point

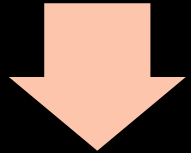
FELDSPAR: REFLECTANCE VS. η (ROUGH GLASS)

Essack et al. (in review)

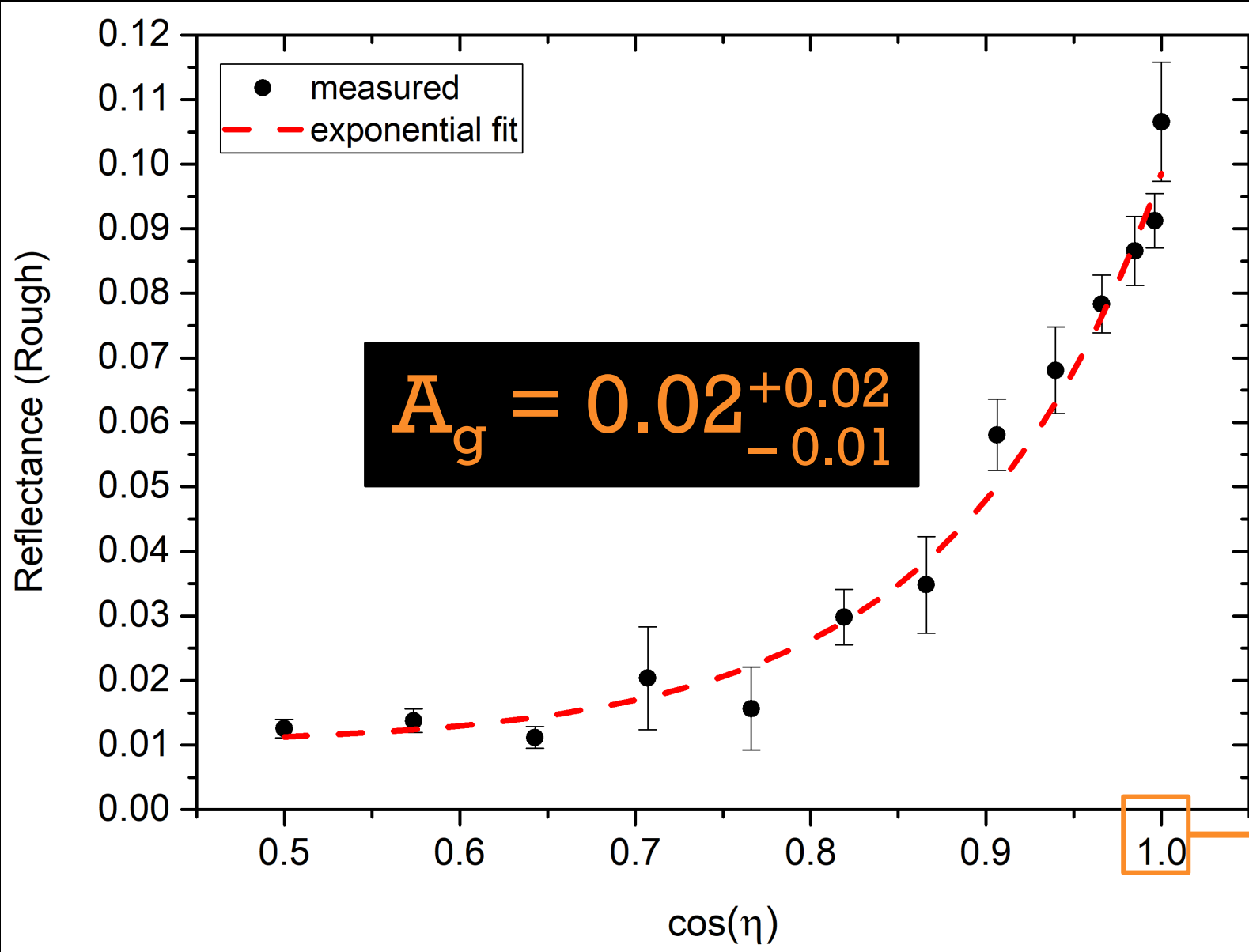
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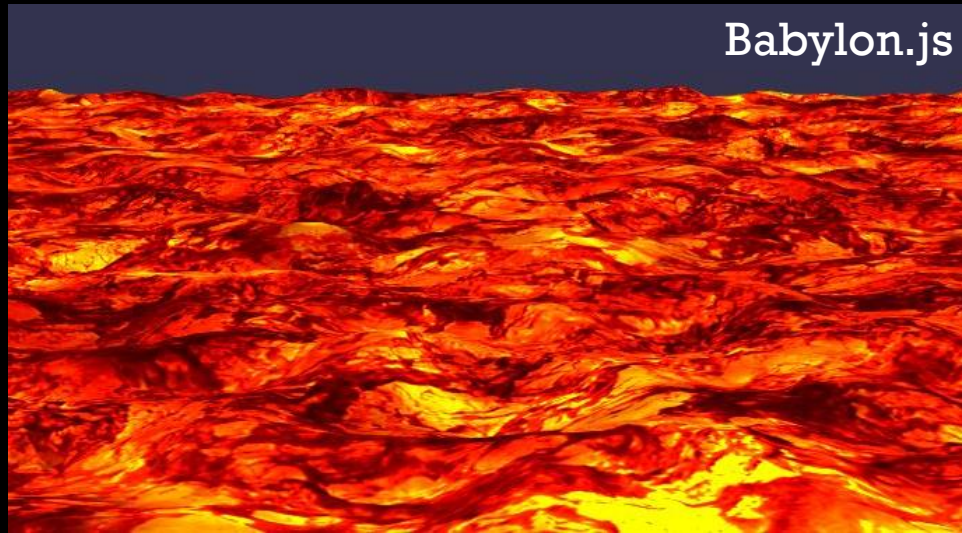
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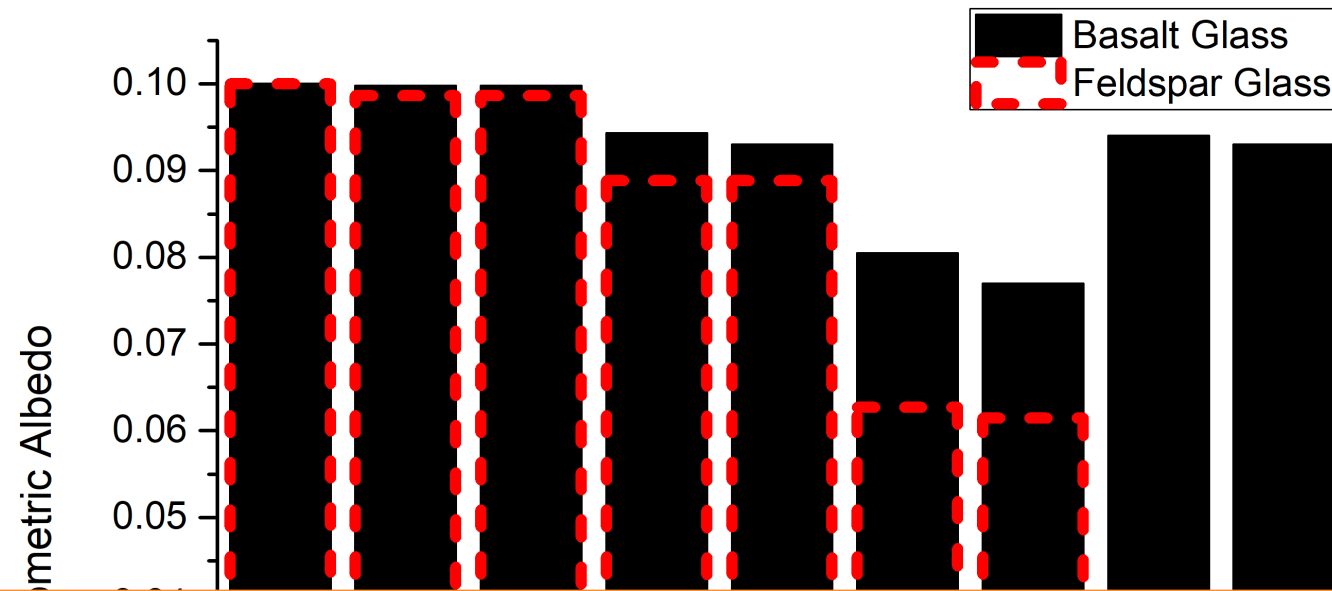
Substellar
point

GEOMETRIC ALBEDO OF A COMBINATION LAVA-GLASS PLANET

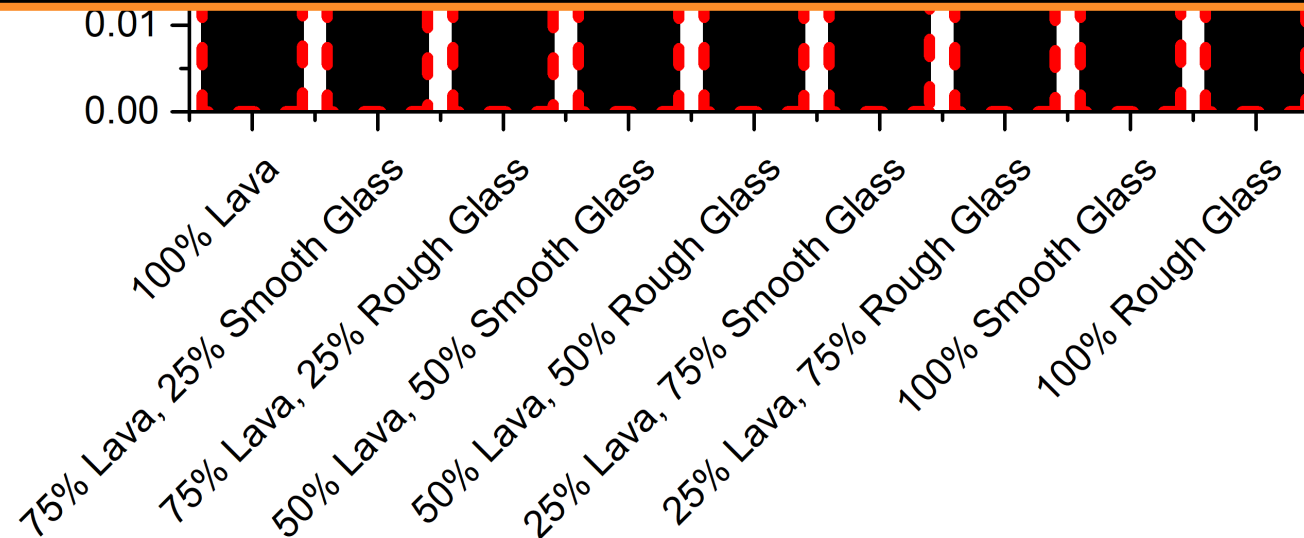


Lava: Specular reflection value
from non-crystalline solids
literature.

Quenched Glass: Reflection
values measured
experimentally.



Reflection from lava and quenched glasses cannot explain the high geometric albedos of hot super Earths.

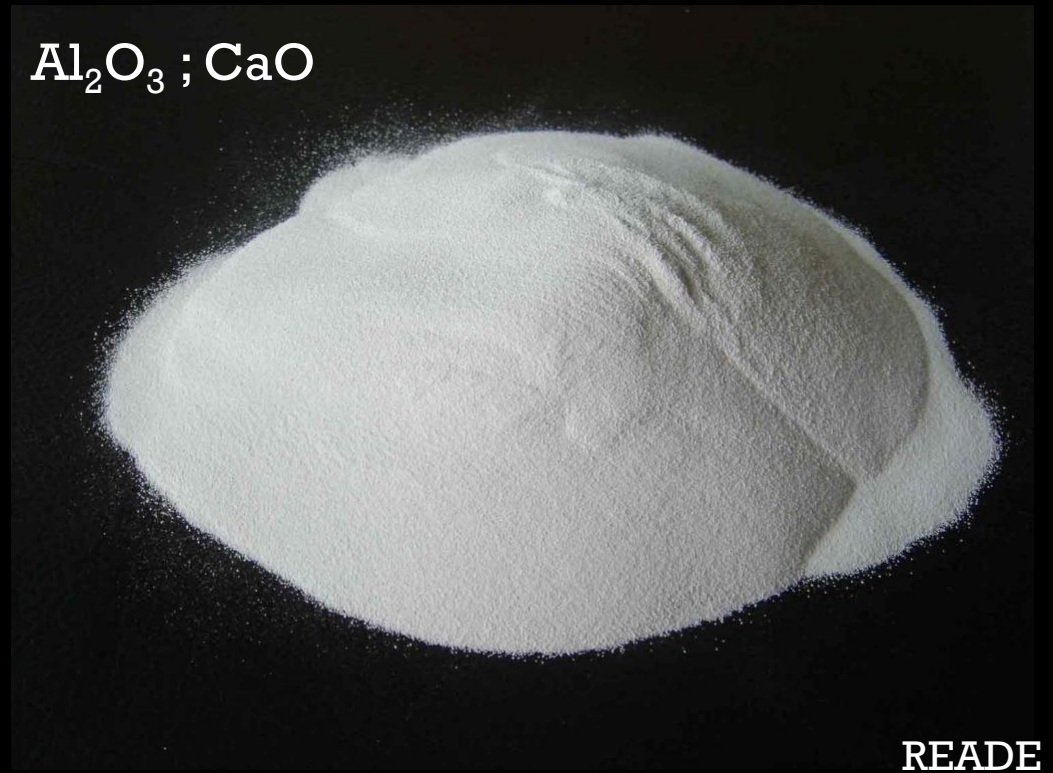


OTHER SOURCES OF REFLECTION

ATMOSPHERES

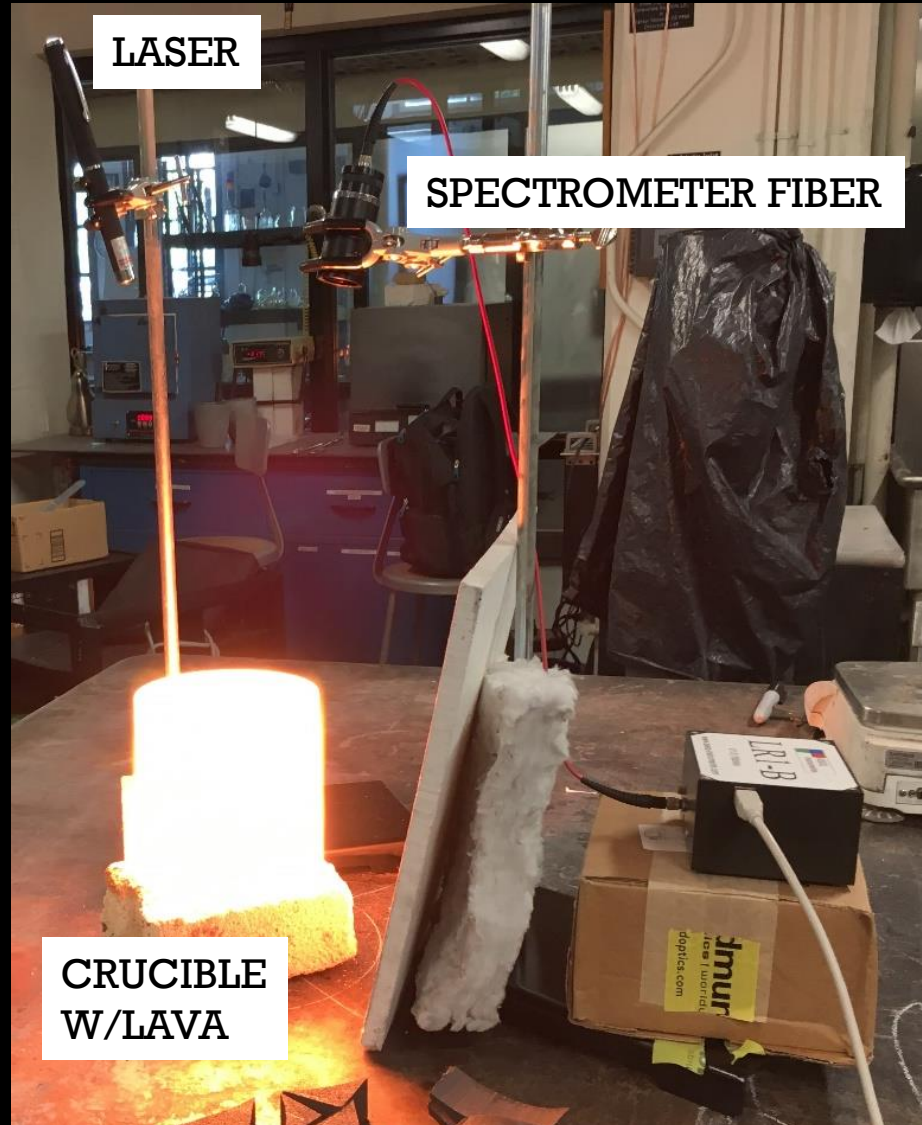


EVOLVED HIGH ALBEDO SURFACES

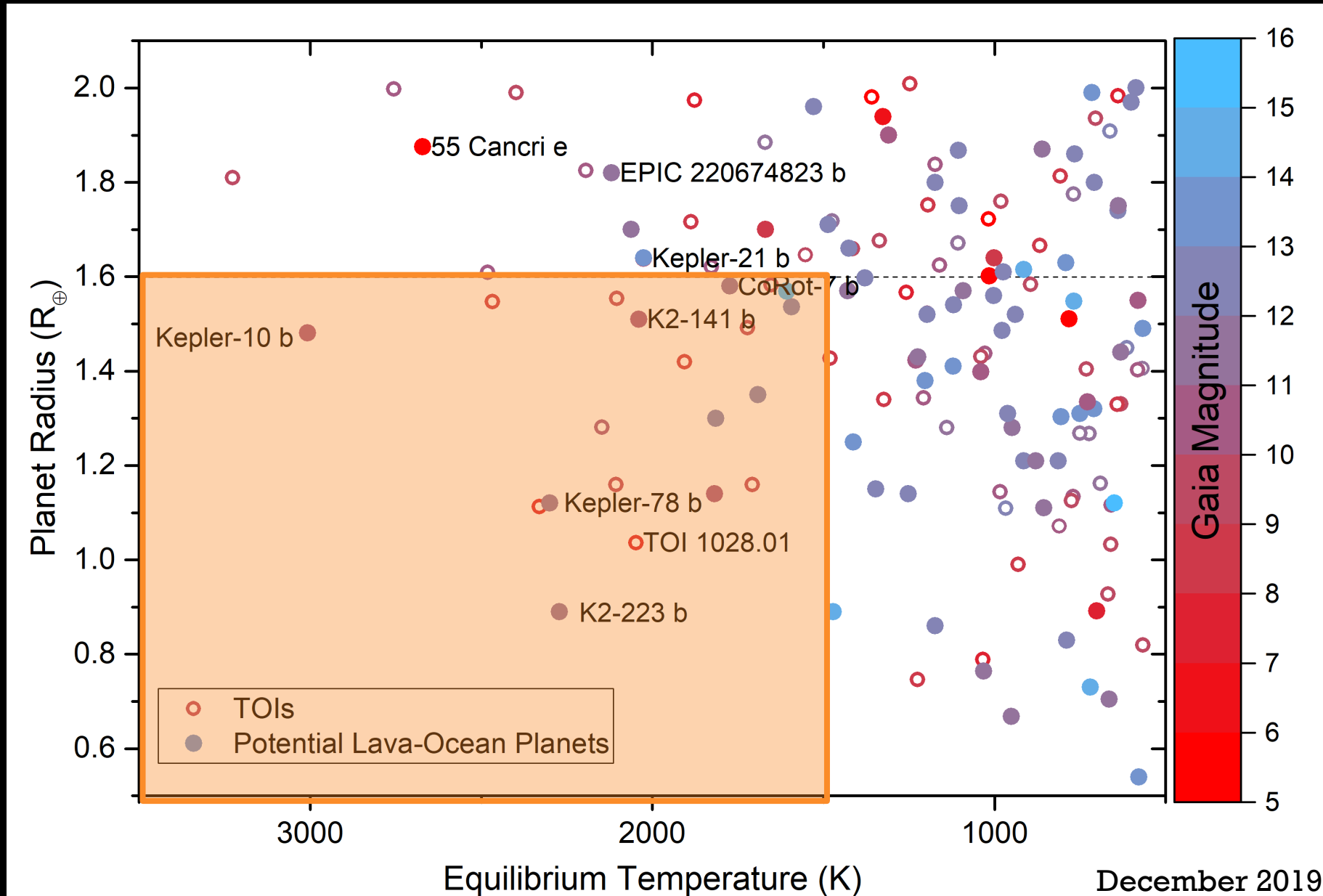


Combining results from Zebger et al. (2005); Hu et al. (2012); Kite et al. (2016).

FUTURE WORK: MEASURING THE ALBEDO OF LAVA



FUTURE WORK: LAVA WORLDS FROM TESS



CONCLUSION

- Lava worlds with solid (quenched glass) or liquid (lava) surfaces have low albedos (< 0.1), and hence a negligible contribution to the high geometric albedos of some hot super Earths.
- The high geometric albedos of hot super Earths are likely explained by atmospheres with reflective clouds or evolved surfaces.
- Validating lava planet candidates from TESS and characterizing them with JWST will allow us to better understand their atmospheres, surfaces, and other properties.