

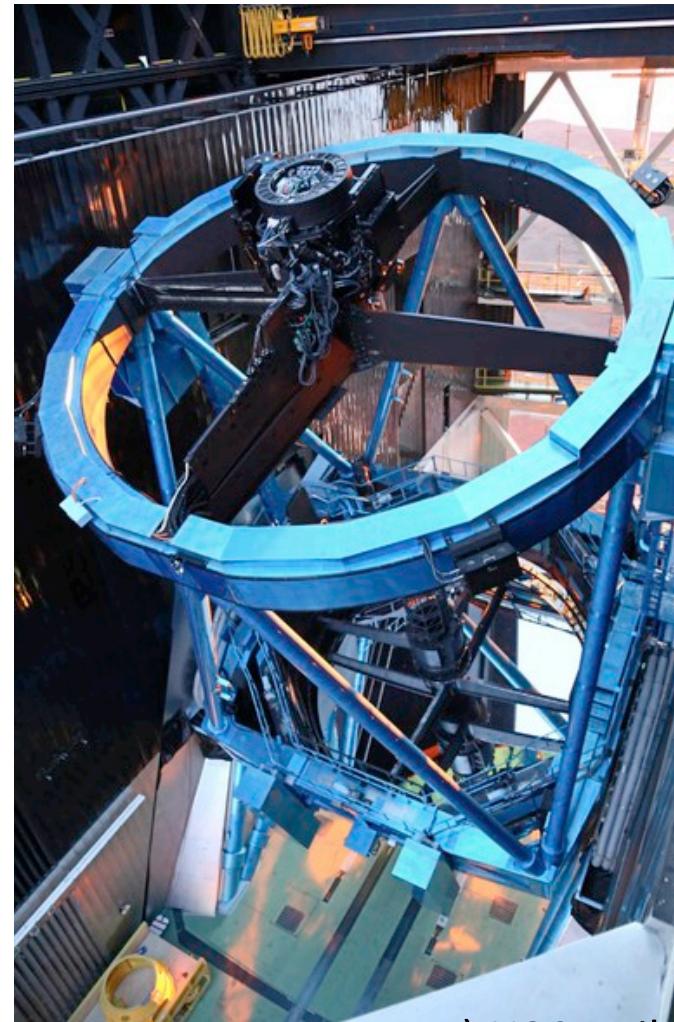
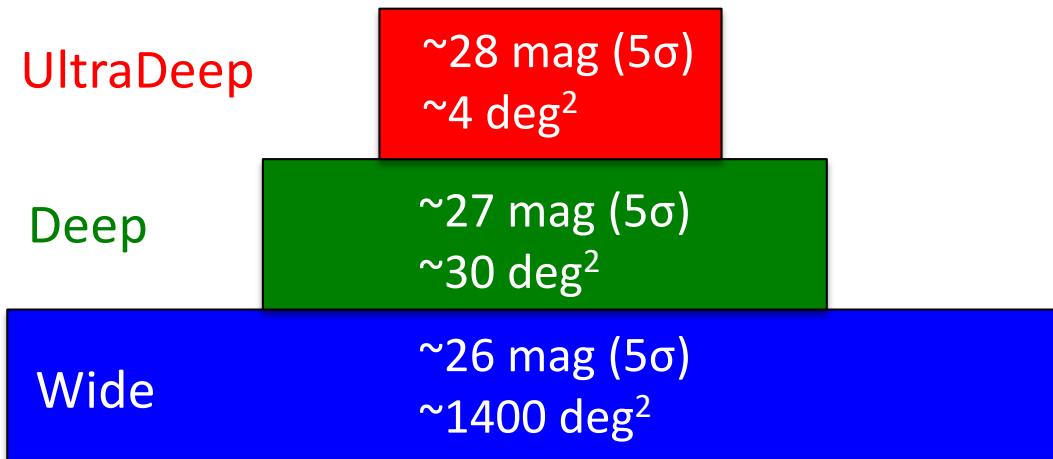
Exploring early galaxies and cosmic structures with Subaru, HST, and ALMA



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MACSJ0717.5+3745
Credit:NASA, ESA and the HST
Frontier Fields team (STScI)

Subaru Hyper Suprime-Cam (HSC) Survey

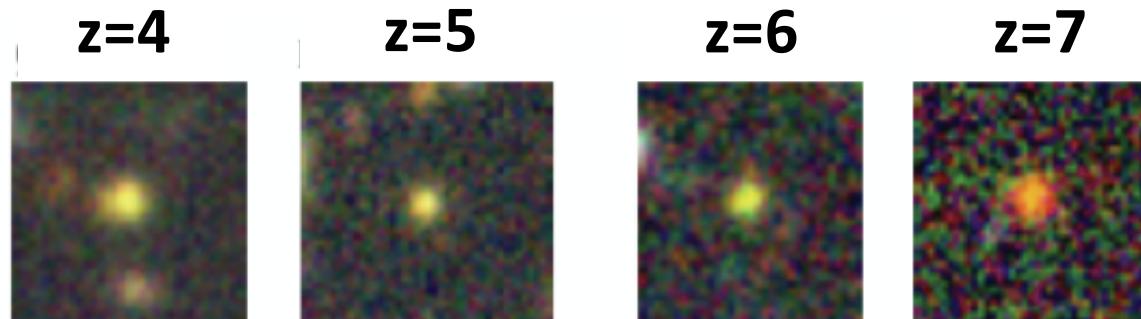


c) HSC Builder's blog

- HSC SSP survey has started since 2014 under the collaboration of Japan, Princeton, and Taiwan.

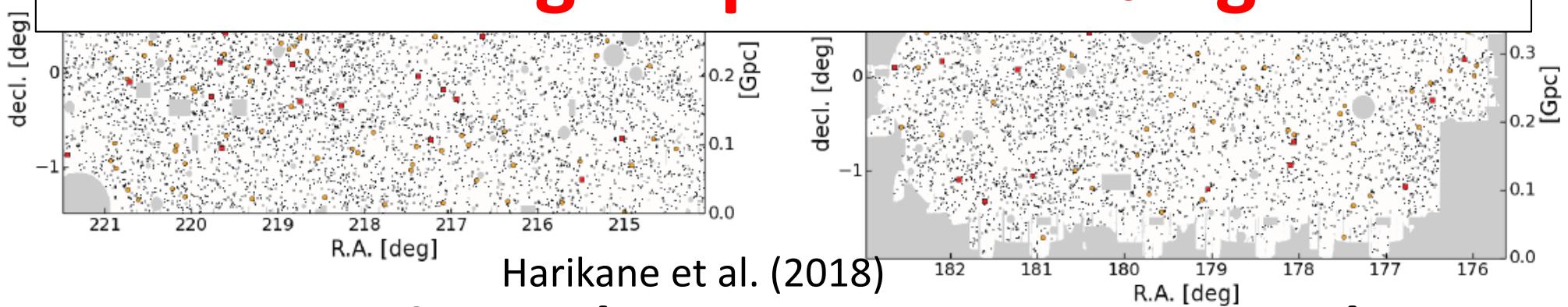
Largest Sample of High-z Galaxies at z=4-7

Examples



Ono et al. (2018)

~100 times larger than prev. samples
First cosmological probe of $z \gtrsim 4$ galaxies



Harikane et al. (2018)

- **579,565 galaxies (+ 2354 Ly α emitters; LAEs)
over 100 deg 2 → 1.4 Gpc 3 (cosmology scale)**

Spectroscopic Follow-up Observations

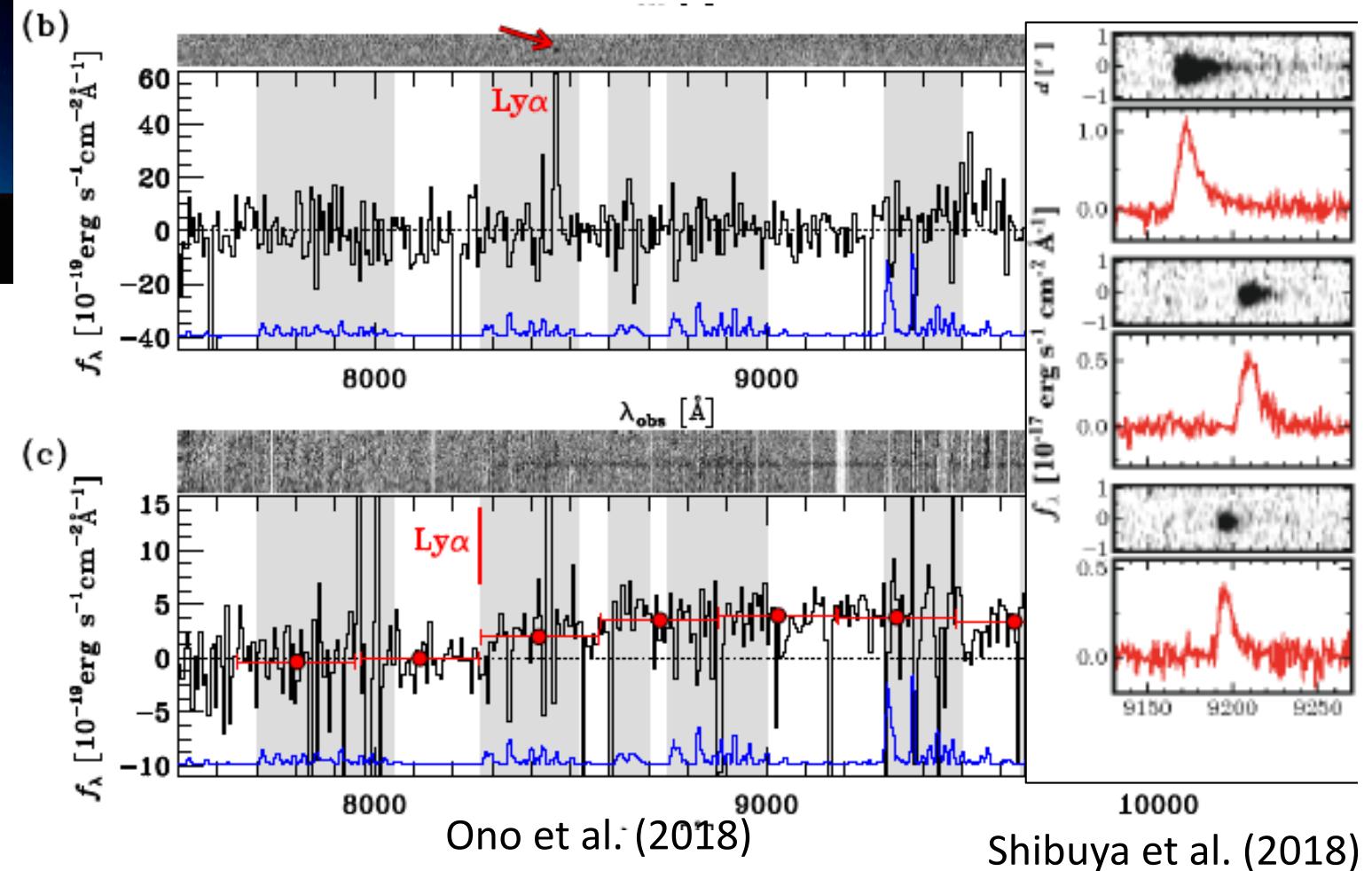
Keck



Subaru



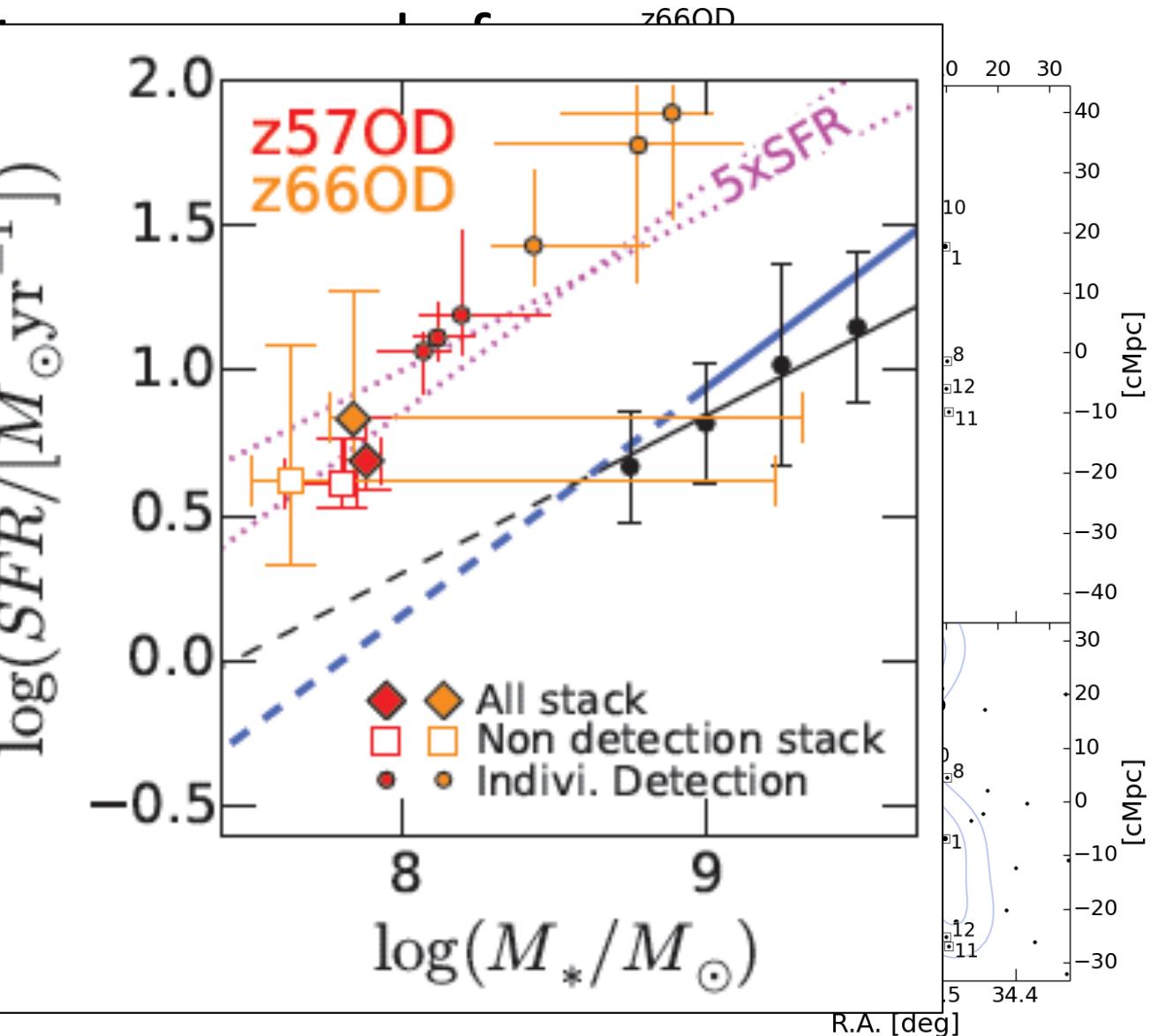
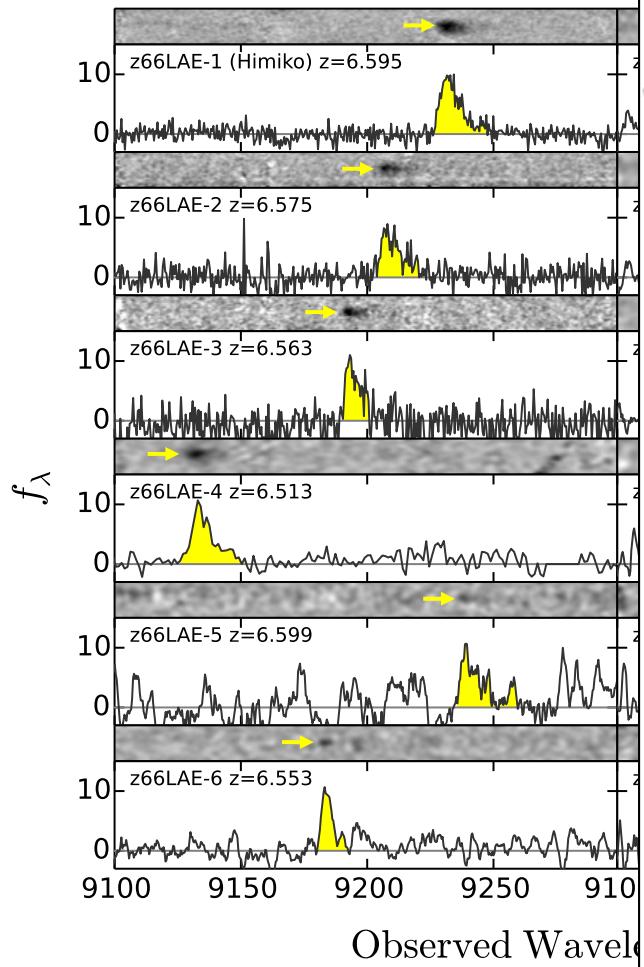
Magellan



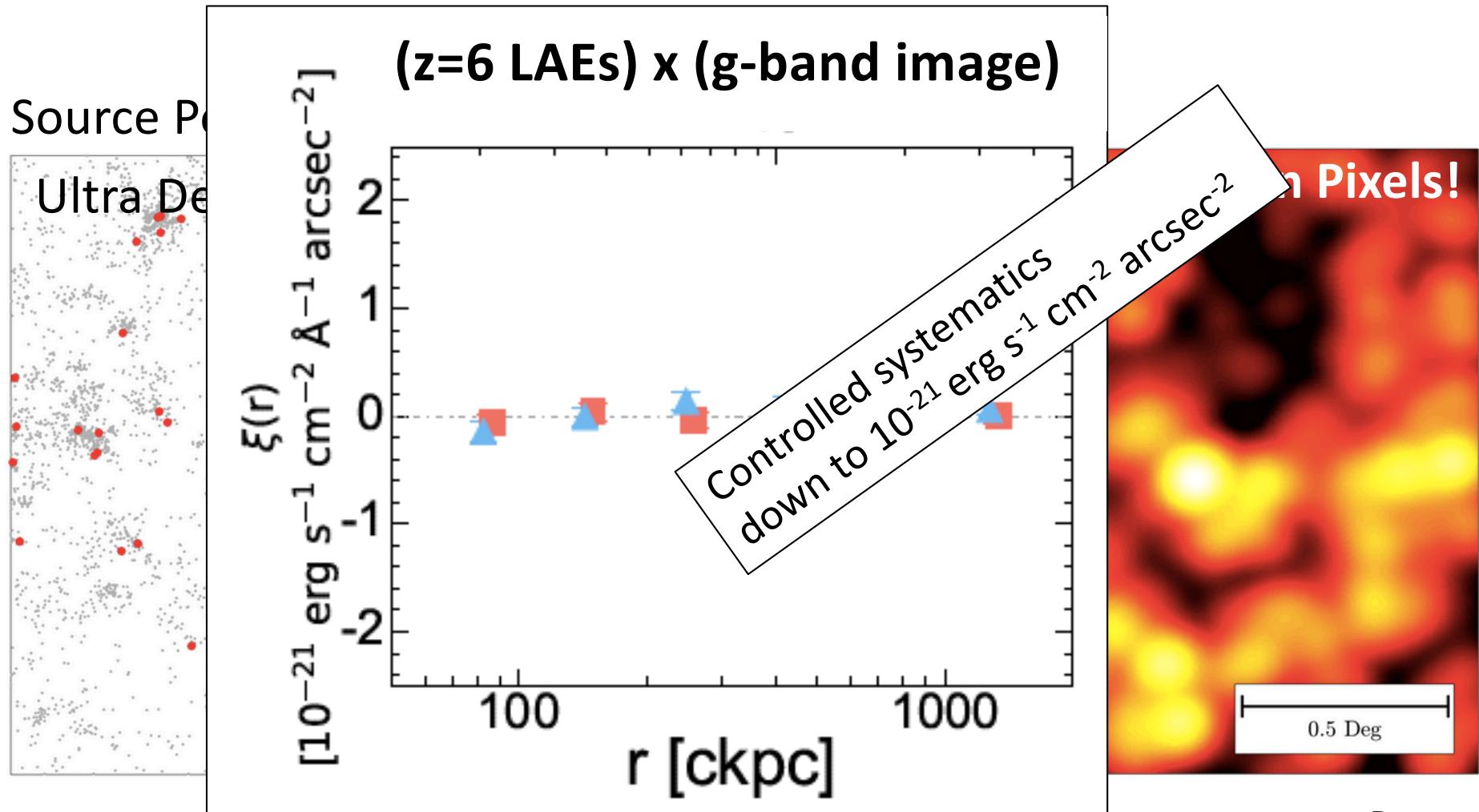
- On-going Keck, Subaru & Magellan spectroscopy → Subaru/PFS spec.
- A total of 348 dropouts at $z=3\text{-}7$ and 174 LAEs at $z=6\text{-}7$ (at the EoR).
- In 1 deg 2 of SXDS LAEs: >80% spec. completeness ($\log L_{\text{Ly}\alpha} > 43.0 \text{ erg s}^{-1}$)

3D Large Scale Structure & Protocluster at $z \sim 7$

- **z660D:** Protocluster
12 galaxies at $z(s)$



Ly α Intensity Mapping Cross-Correlation with the LAEs

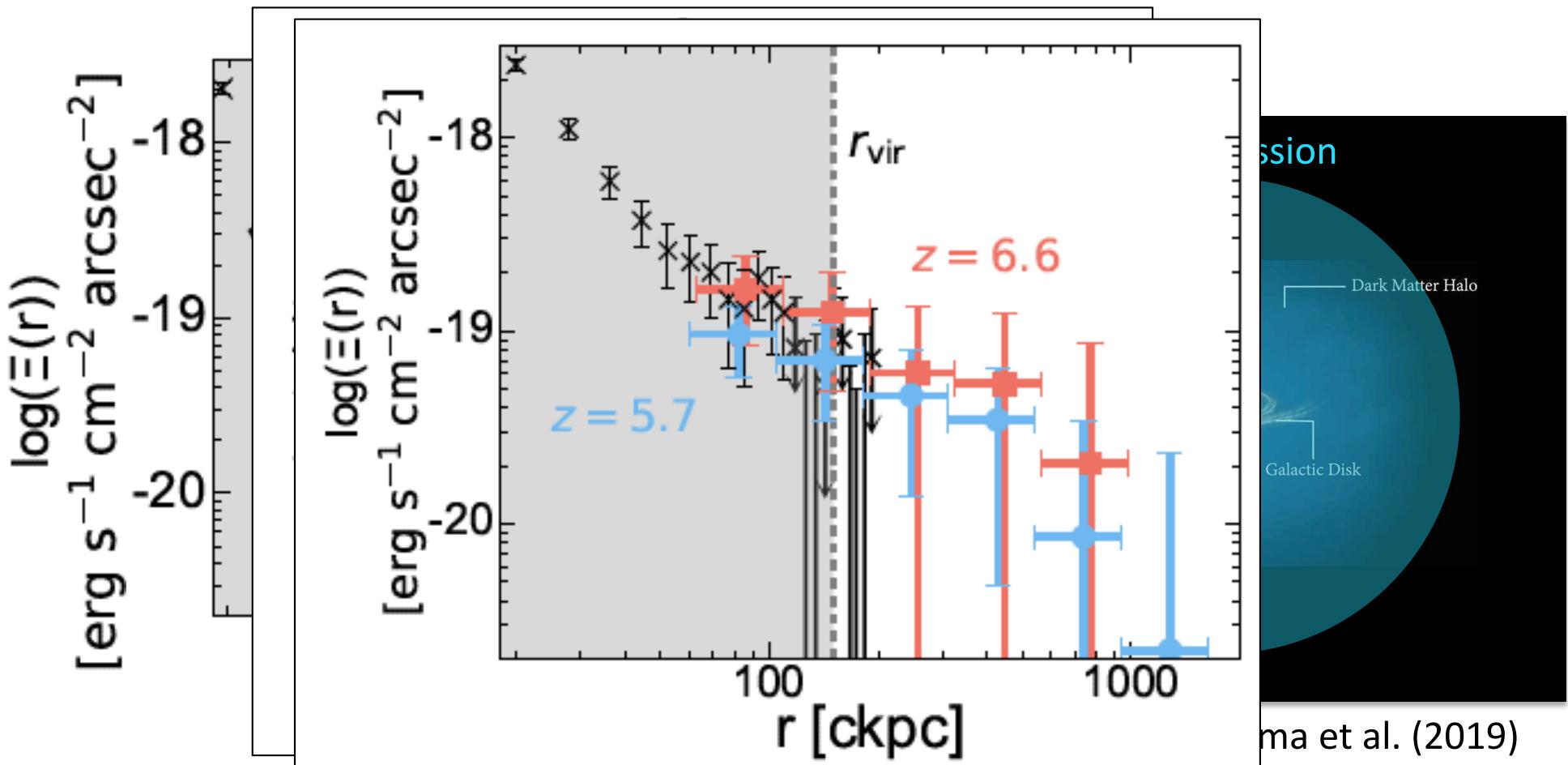


- Careful evaluation of systematics w uncorrelated sources for calibration (Kakuma et al. 2019)

Breysse+16

Ly α Intensity Mapping

Cross-Correlation with the LAEs



- HSC: Largely extended Ly α emission with a size of 200-1000 comoving kpc (about 5 times larger than virial radius of dark matter halo).
 - Physical origin? Ly α resonance of associated dwarf galaxies/cold accretion?

Reionization effects? -> Comparison with $z=5.7$ and 6.6 results

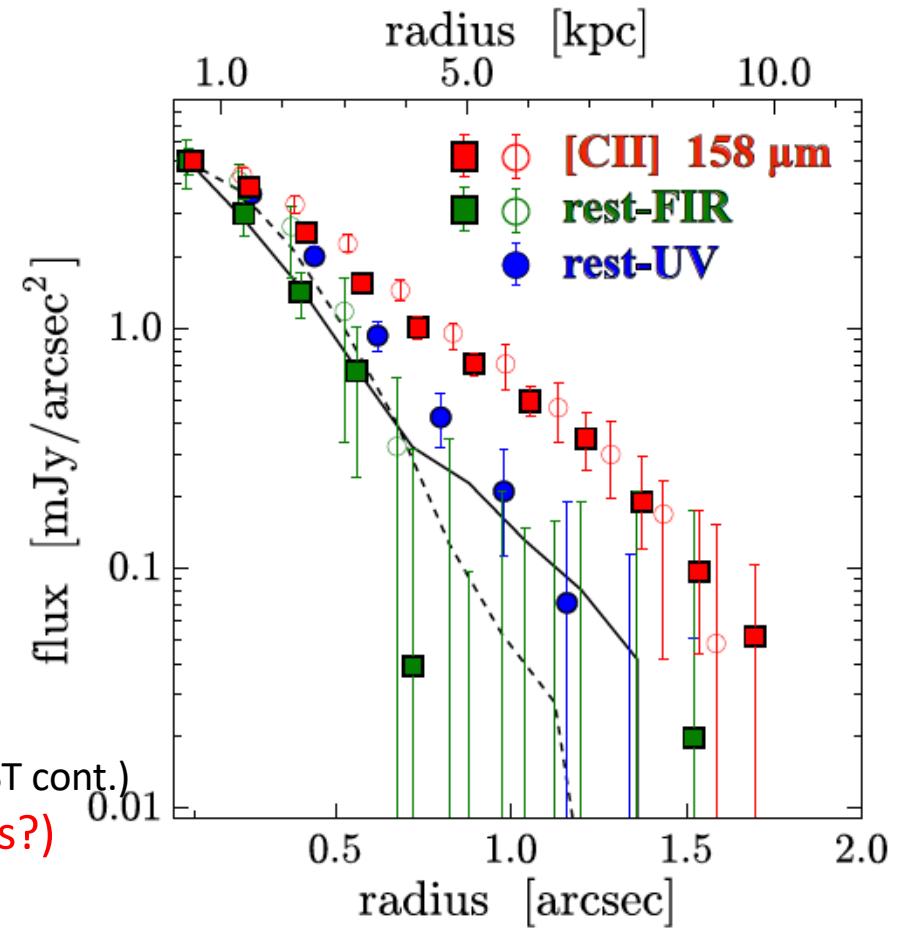
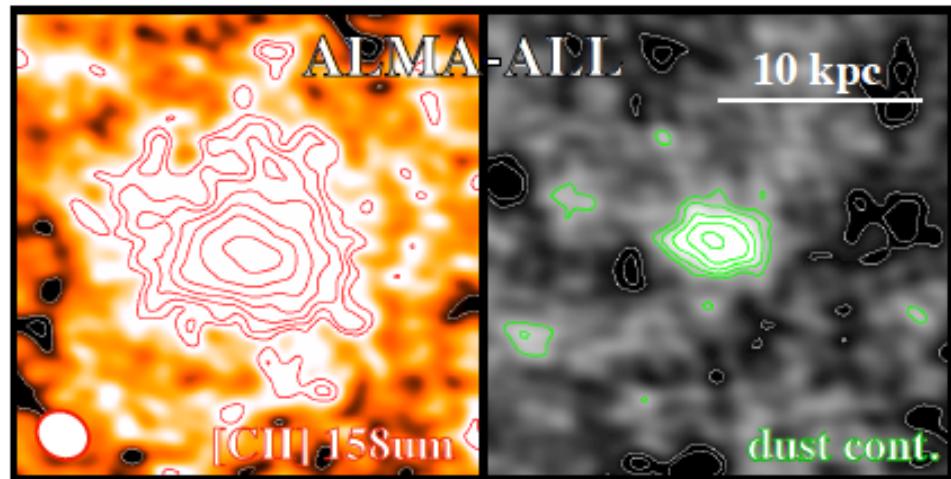
Ma et al. (2019)

Metal Halo Suggested by ALMA obs

Fujimoto et al. (2019)

FIRST IDENTIFICATION OF 10-kpc SCALE [CII] 158 μ m HALOS AROUND STAR-FORMING GALAXIES AT $z = 5 - 7$

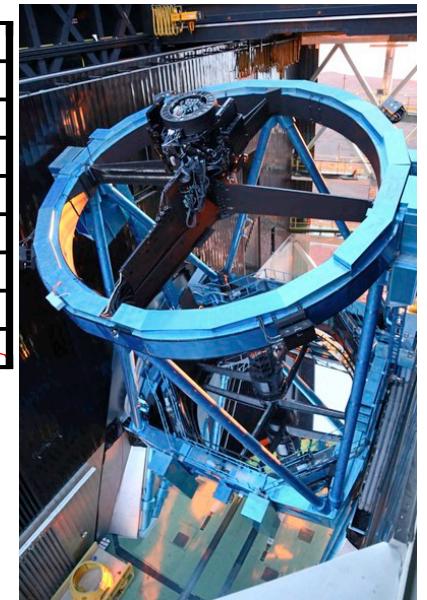
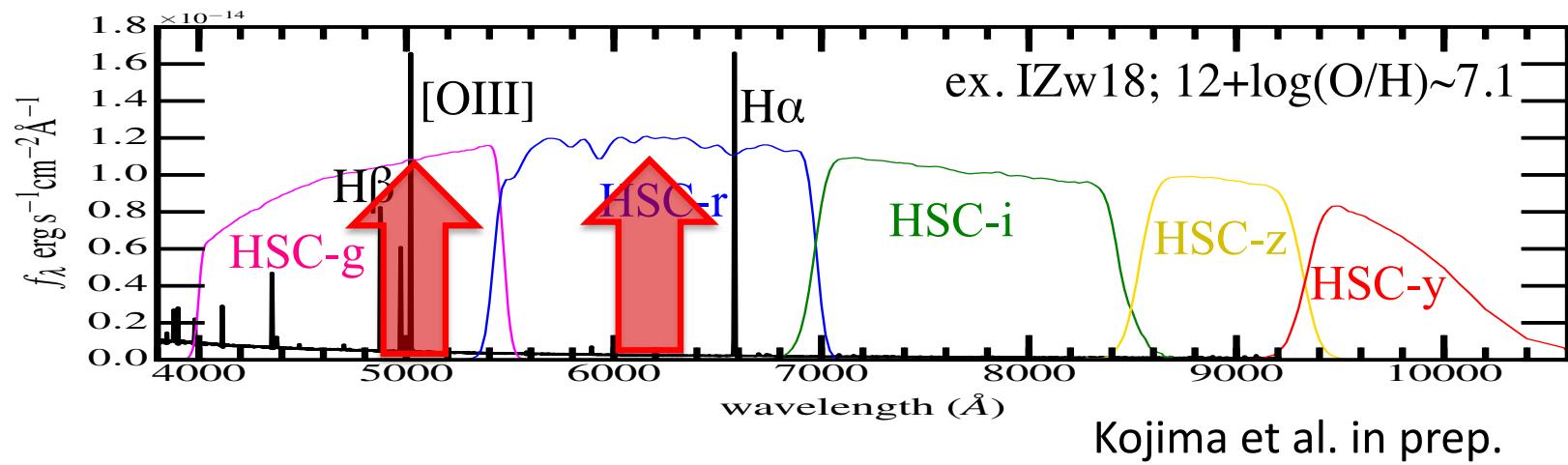
SEIJI FUJIMOTO¹, MASAMI OUCHI^{1,2}, ANDREA FERRARA^{3,4}, ANDREA PALLOTTINI³,
R. J. IVisON^{5,6}, CHRISTOPH BEHRENS³, SIMONA GALLERANI³,



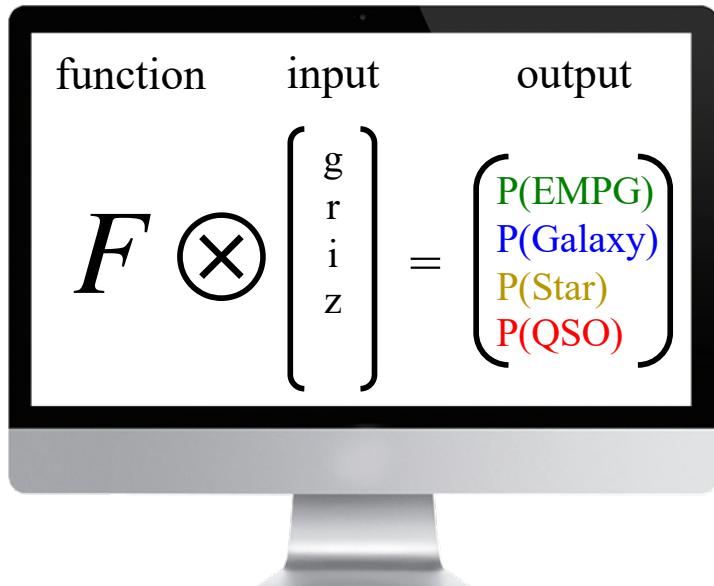
- 10-kpc [CII]158 μ m halo at $z=5.153-7.142$
 - C⁺ gas extends 5x more than SFR regions (dust/HST cont.)
 - Interestingly, comparable with Ly α halo (HI gas?)
 - Simulations do not explain the halo
- important mechanism of outflow? (e.g. cold-mode)

Early Galaxy ISM?

Extremely Metal Poor Galaxy at z=0



+ Machine learning technique



SED models ($30,000 \times 4$)

Beagle

Chevallard & Charlot 2016

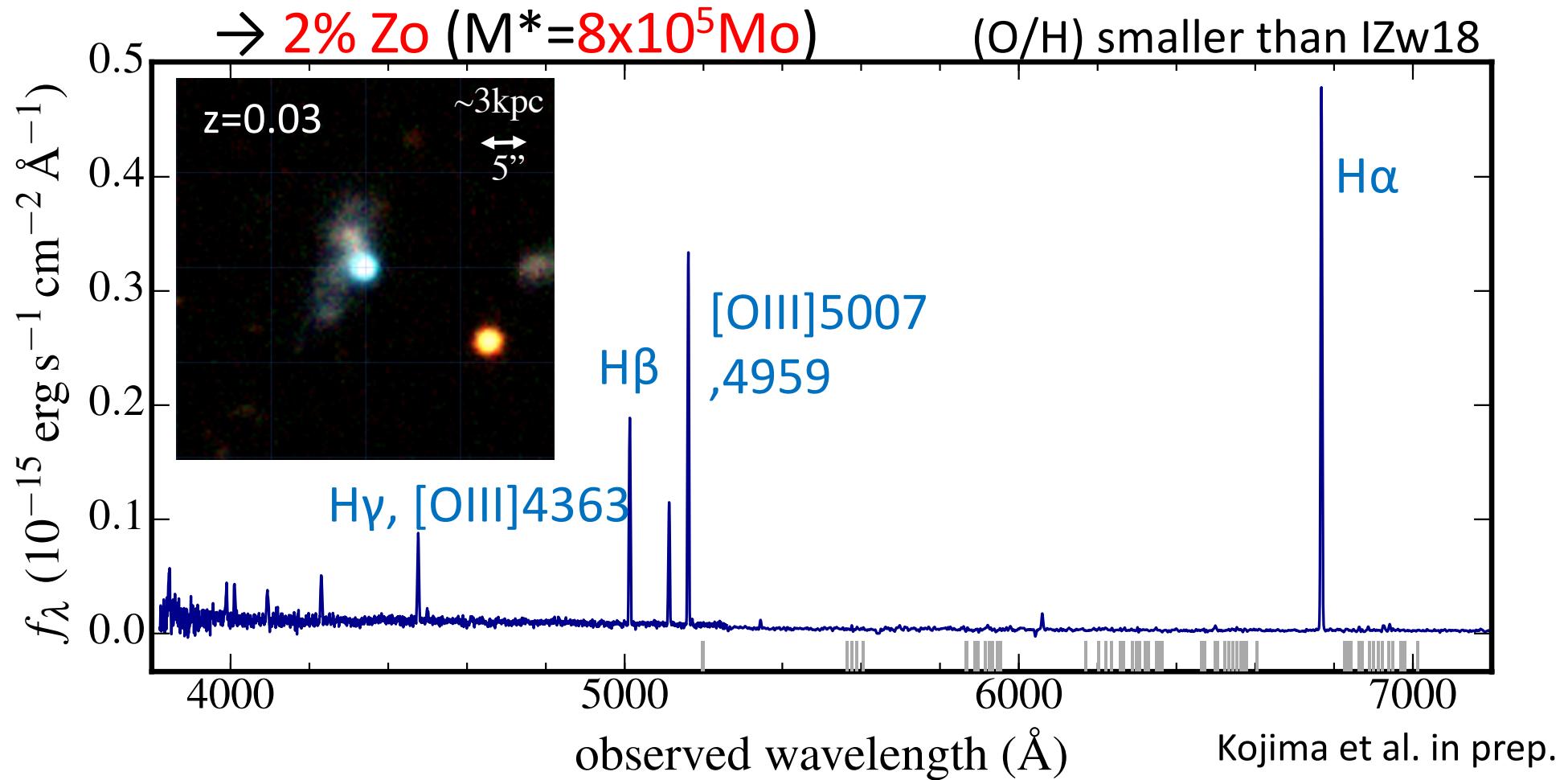
Stellar model

Castelli & Kurucz 2004

QSO composite Selsing et al. 2016

Subaru/HSC: 10 Metal Poor Galaxies

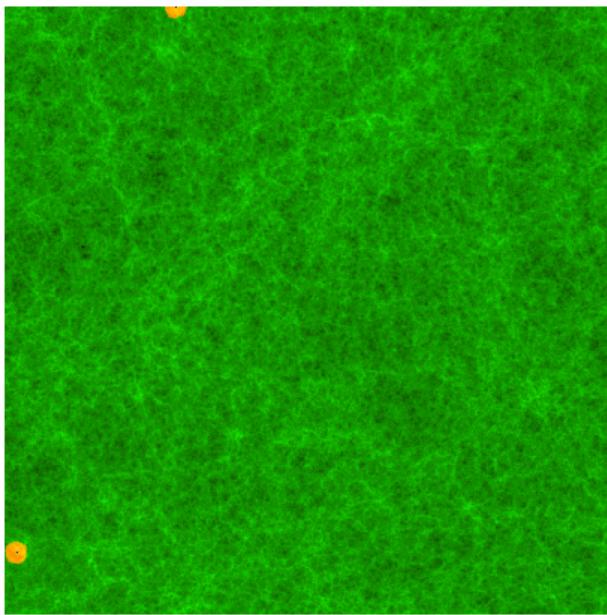
Example: HSC J1631+4426



- Paper I. : Subaru/HSC+ML survey design (Kojima et al.)
- Paper II. : Gas densities and ionization states (Kojima et al.)
- Paper III.: Morphological properties (Isobe et al.)

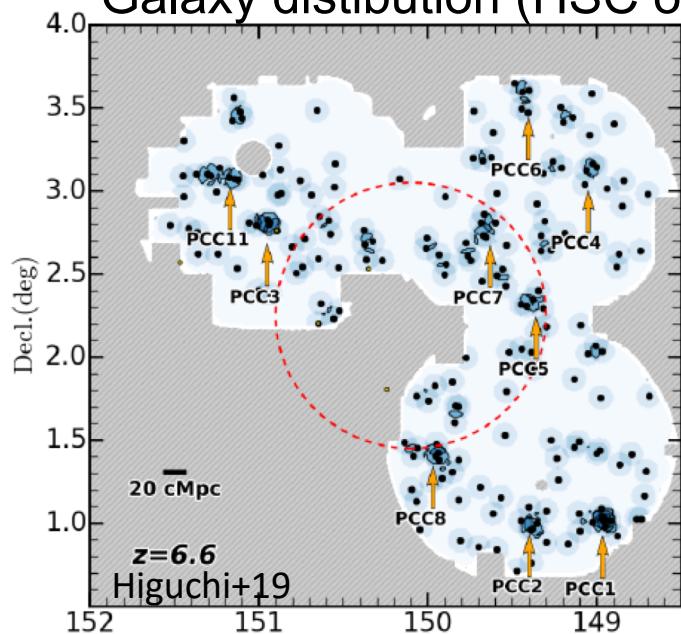
Future Prospects: Galaxy-21cm Cross Correlation

RT simulations

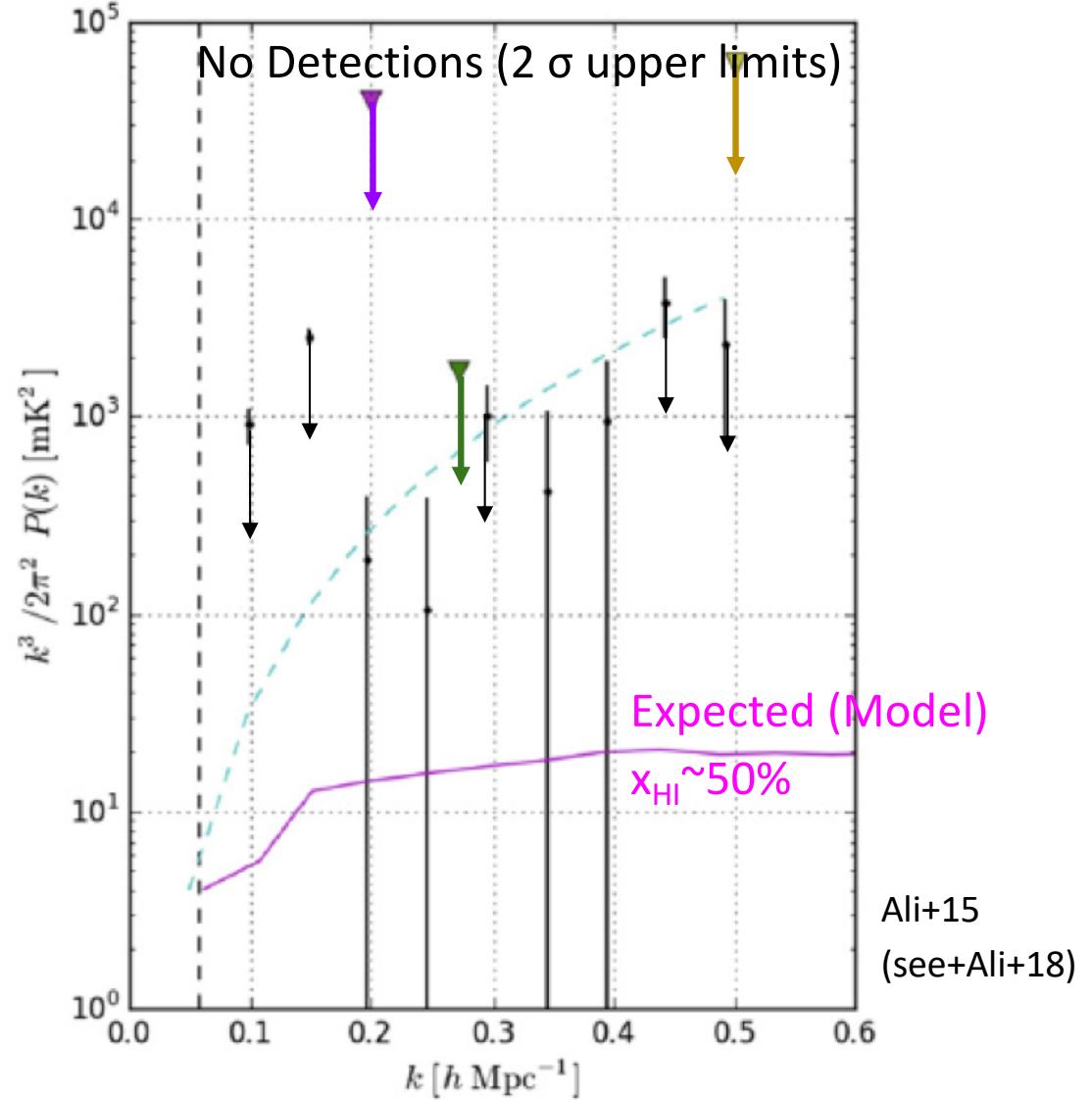


Galaxies (blue dots),ionized bubbles (orange) (Iliev+06)

Galaxy distribution (HSC obs)



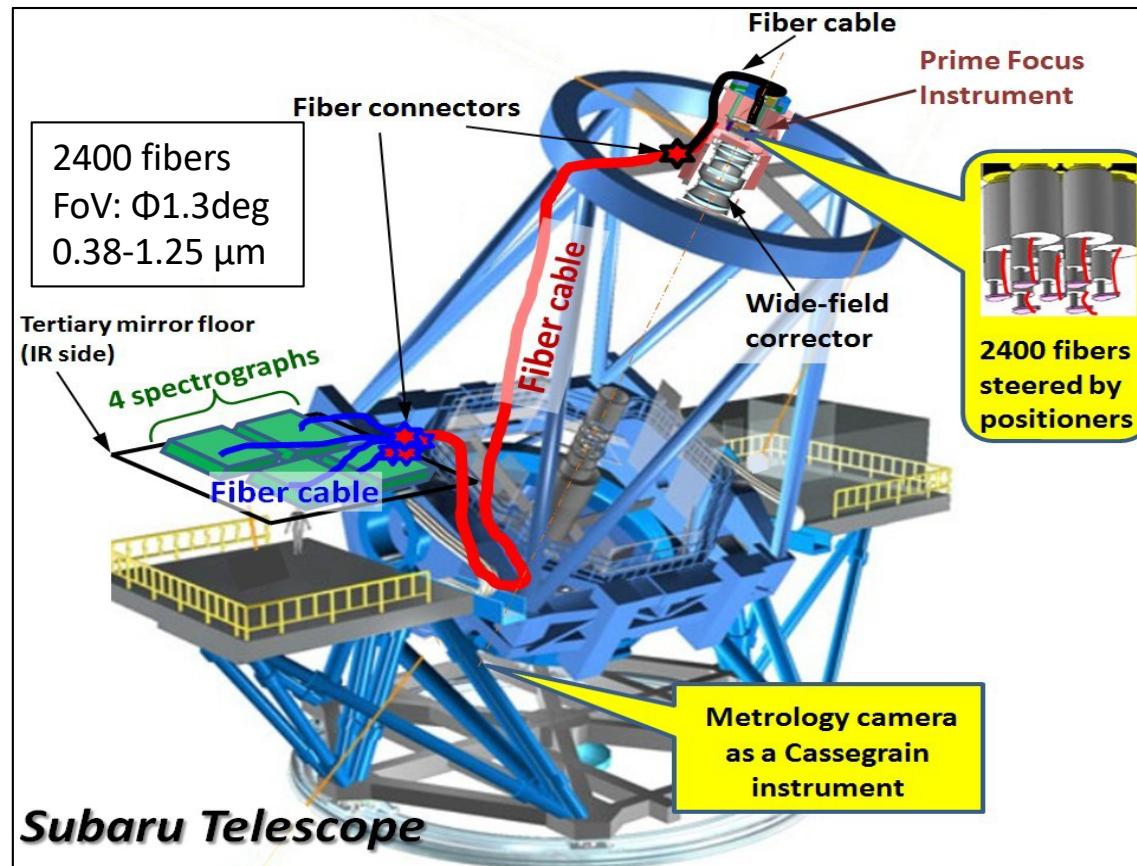
Epoch of Reionization (EoR) HI 21cm Auto Power Spectrum



No detection of 21cm EoR signals, due to FG sys (Gehlot+18)

- Cross Correlation with real signal of galaxies at EoR

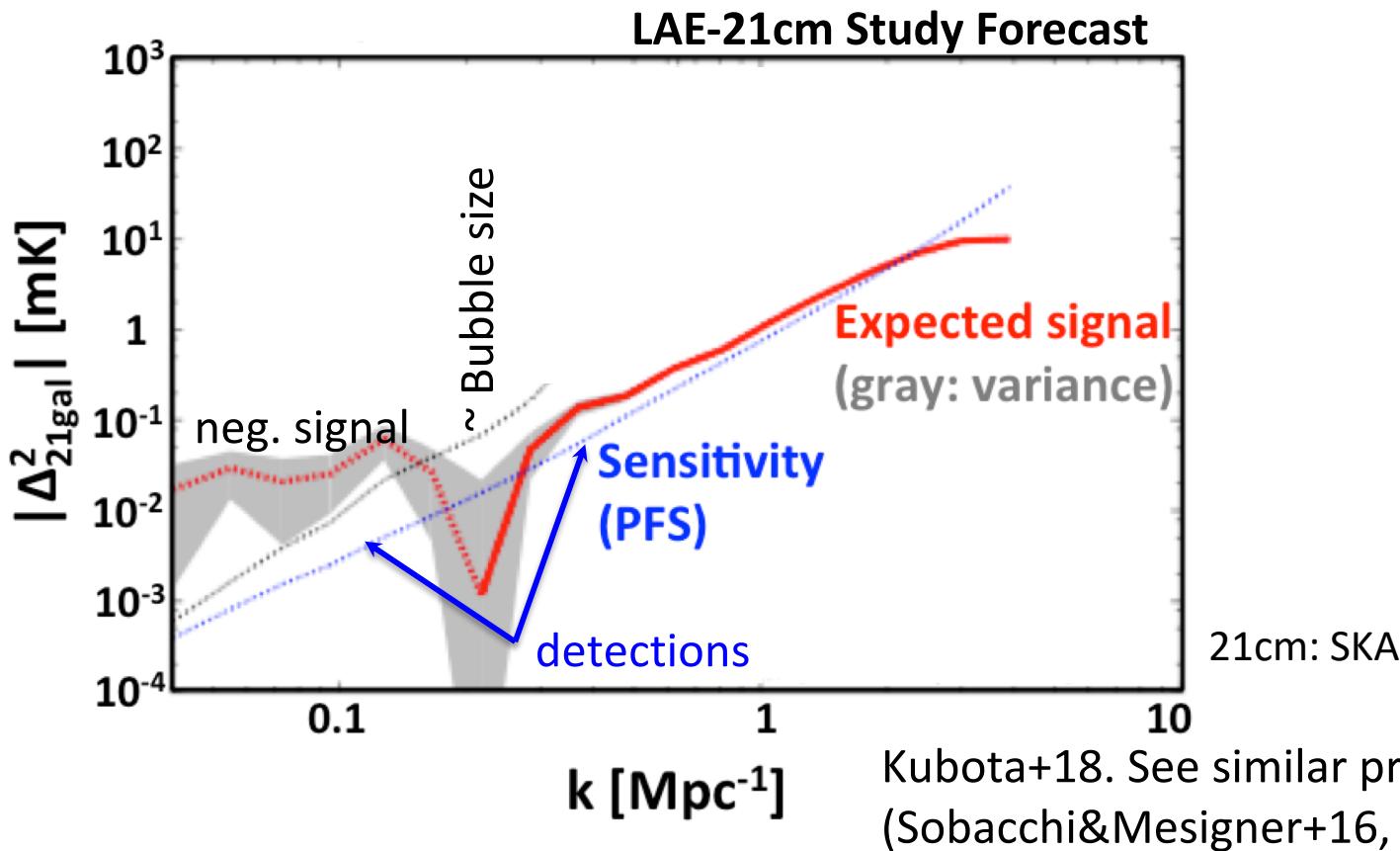
Subaru Prime Focus Spectrograph (PFS)



- Under the collaborations with many institutes over world (planned FL 2021)
- Spectroscopy for \sim 10,000 LAEs at the EoR ($z=6-7$) over 15 deg^2 area

Galaxies (PFS) and HI 21cm (MWA, SKA) \rightarrow Cross-correlation signals

First Detection of EoR HI 21cm Signal by Galaxy-21cm Cross Correlation



- Goal-1: Detection of the cross-correlation signals -> Evidence of early cosmic HI struc.
 - Positive cross-correlation at $k \sim 0.4 \text{ Mpc}^{-1}$ at ~ 5 sigma
 - Negative cross-correlation at $k \sim 0.1 \text{ Mpc}^{-1}$ at ~ 3 sigma
- Goal-2: Determination of the transition scale at $z=6.6$ with $\Delta k \sim 0.1 \text{ Mpc}^{-1}$ accuracy
 - First definitive evidence of cosmic ionized bubbles

Summary

- Recent studies of early galaxies and reionization
 - Subaru HSC survey (579,565 gals at $z=4-7$ in 100 deg^2)
 - 1) 3D galaxy map at $z=6.6 \rightarrow$ proto-cluster at $z=6.585$
 - 2) Detection of Ly α emission beyond r_{vir} , 200-1000 ckpc
 - 3) ALMA identification of extended [CII] halo (10 pkpc)
 - 4) Extremely metal poor galaxies at $z=0$ incl. 2% Z_0
- Subaru/HSC+PFS survey and 21cm observations
 - Identifying EoR 21cm signals and ionized bubbles